

He discharges with transition on GOLEM - spectroscopic study

Beatriz García¹

Supervisors: Diana Naydenkova³, Vladimir Weinzettl³
Helpers: GOLEM team²

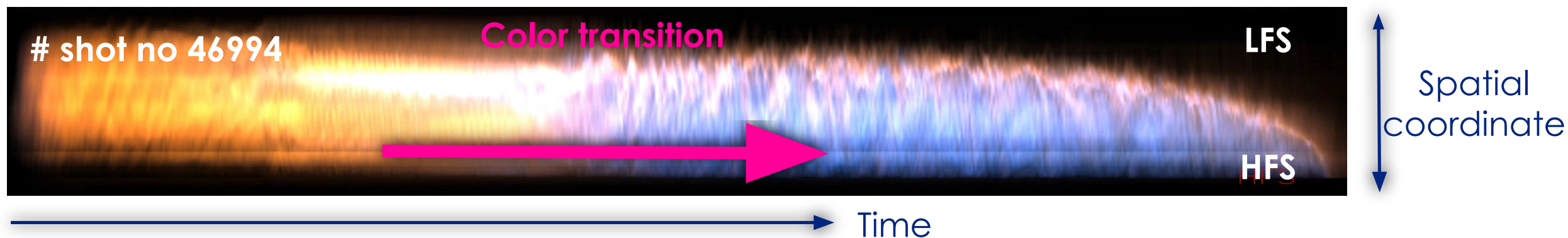
Affiliations:

¹ Aix-Marseille University, ² FNSPE, Czech Technical University in Prague , ³ Institute of Plasma Physics of the CAS



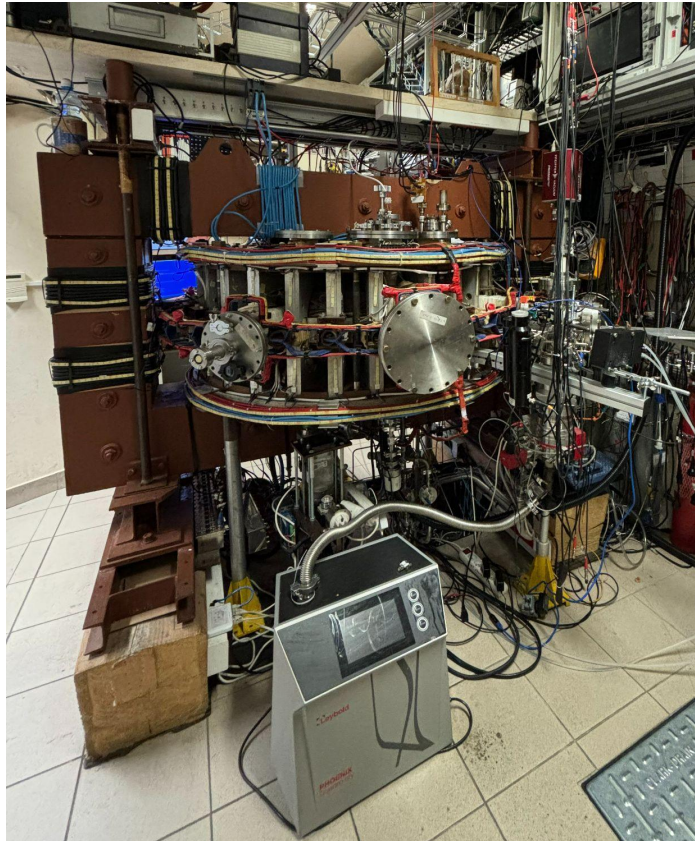
Motivation

What is behind the transition in He discharges in GOLEM?



High-speed camera (40 kfps, time evolution of the narrow spatial stripe)

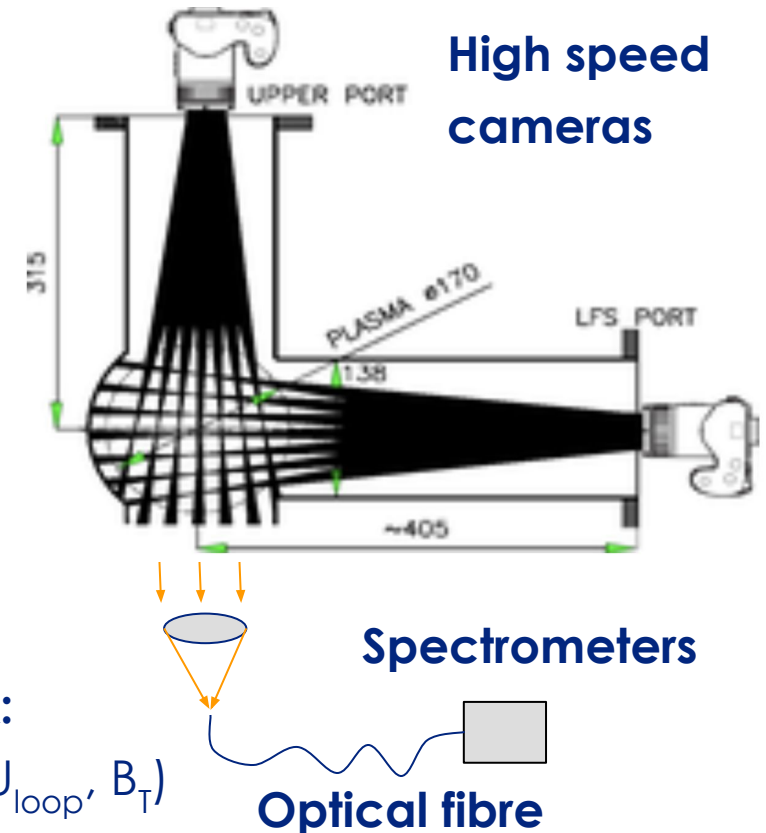
- Small university tokamak at FNSPE CTU in Prague



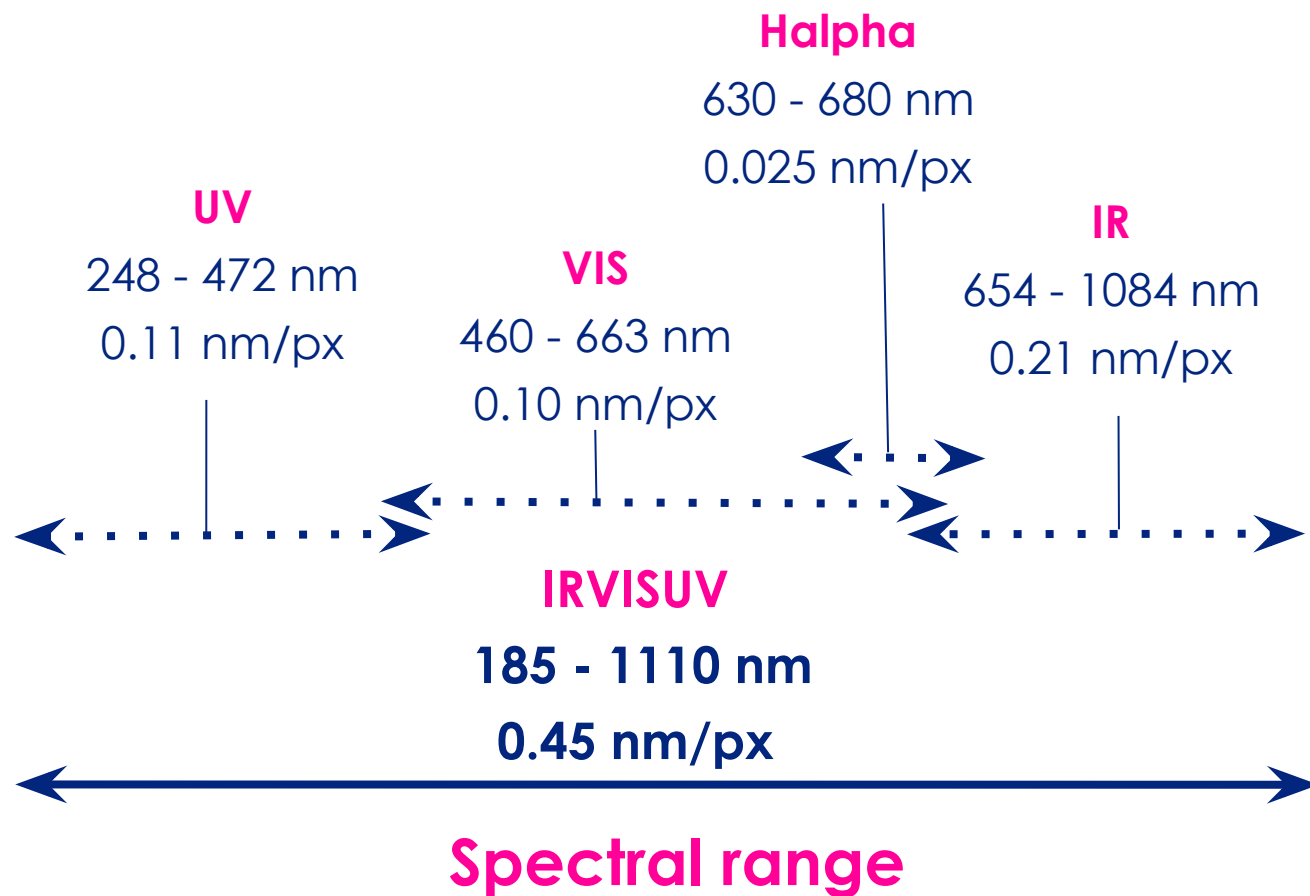
I_p	$< 5 \text{ kA}$
B_T	$< 0.5 \text{ T}$ (sinusoidal)
$\langle n_e \rangle$	$< 3 \times 10^{19} \text{ m}^{-3}$
$T_e(0)$	$< 100 \text{ eV}$

Plasma diagnostics used for the task:

- Basic magnetic diagnostics (I_p , U_{loop} , B_T)
- High speed cameras
- Compact spectrometers (NUV-NIR spectral range)



Compact spectrometers (Czerny-Turner type with grating & CCD)



Spectrometers at GOLEM connected with the tokamak via optical fibers

Previous Work

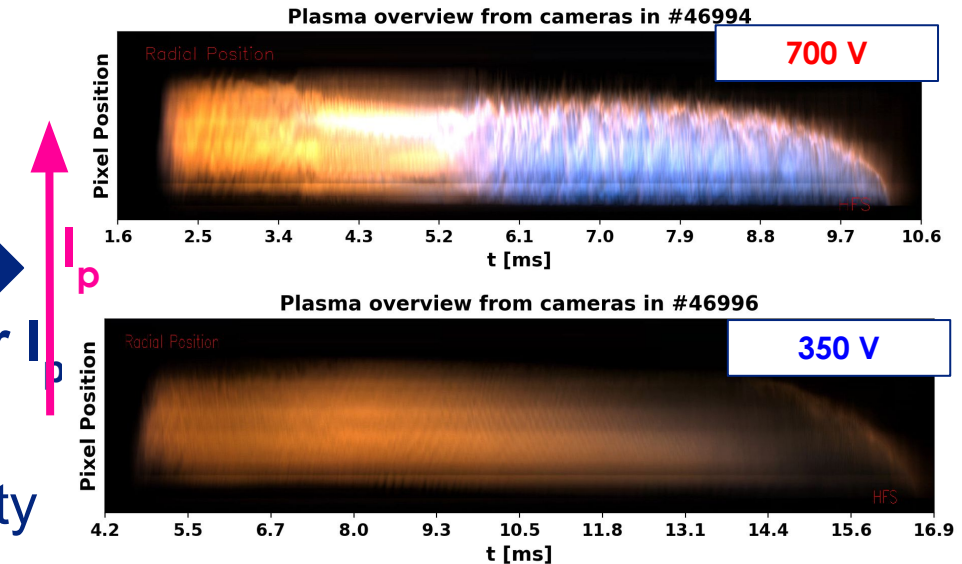
EMTRAIC 2024

Beatriz García, Momoko Kato

Supervisors: Vladimir Weinzettl³, Diana Naydenkova³
Helpers: GOLEM team²

EMTRAIC 2024

- 1) I_p Scan → Smooth behaviour in I_p ($I_{p,crit}$ not found)
→ Color transition more intense for higher I_p
- 2) Effect of breakdown and B_T → no B_T dependence.
→ changes MHD activity
- 3) Spectra analysis (IRVISUV) → Light ($Z < 20$) impurities found
to be responsible for color transition 😊
→ Lines of higher ionization states not confirmed
(but not ruled out either)



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H I, O I, O II, N II, C II, Cl II

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(but not ruled out either)  **He II, O III, N III, C III, Cl III?**



Research Internship

August 2025

Beatriz García

Supervisor: Diana Naydenkova³
Helpers: GOLEM team²

GOAL

→ Search for higher ionized ions: He II, C III, Cl III, O III, N III

HOW?

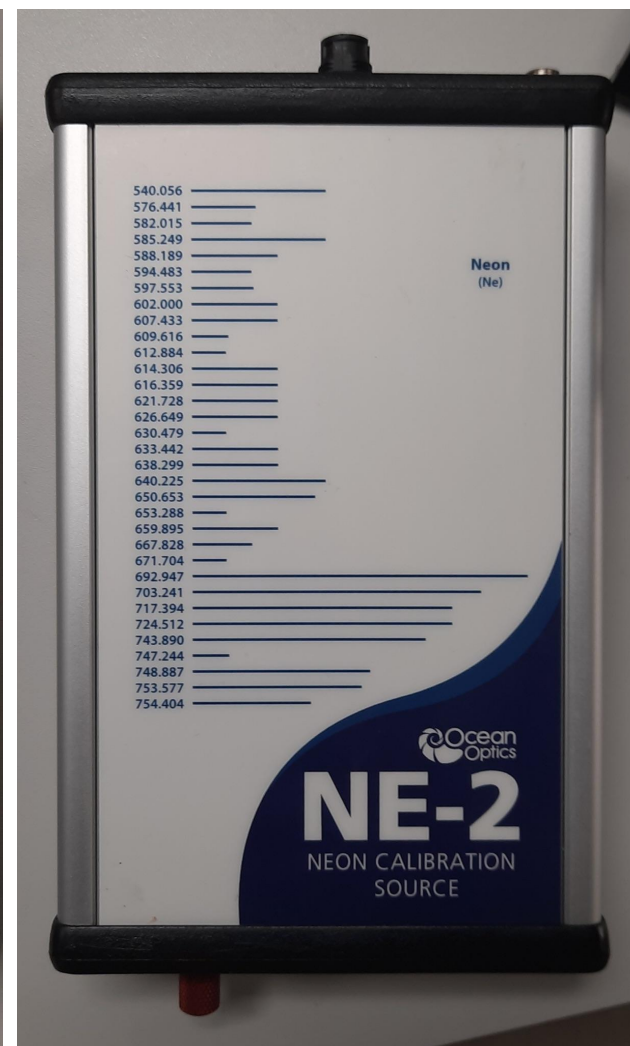
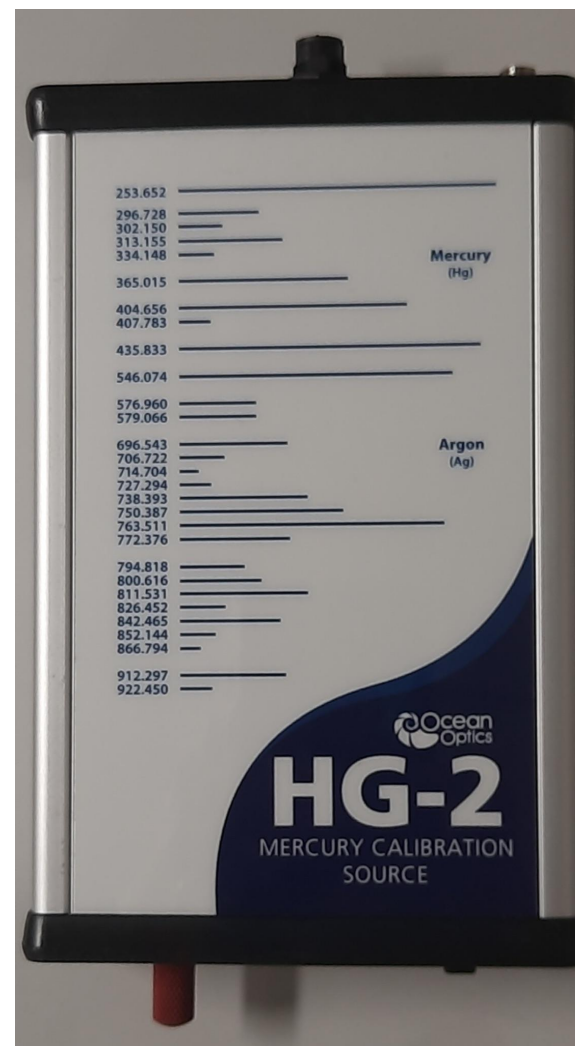
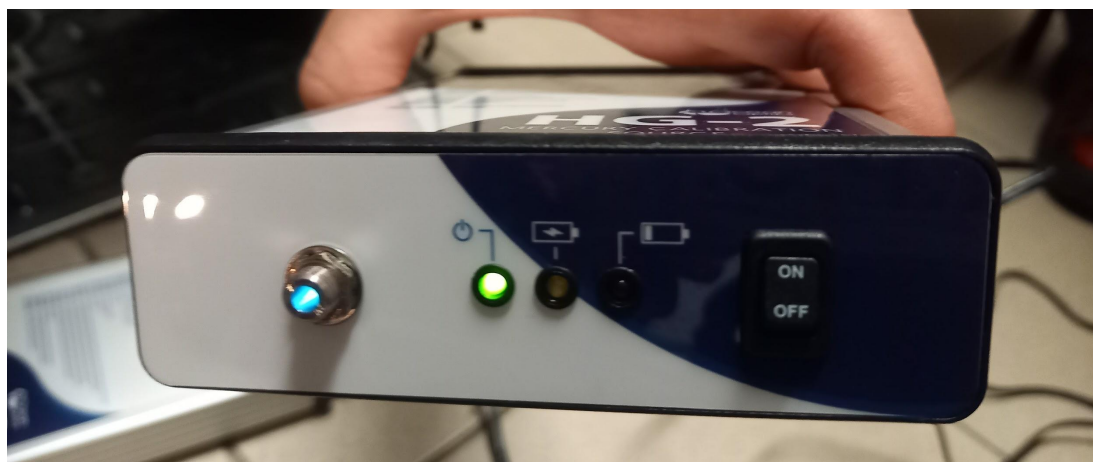
→ Use Higher Resolution Spectrometers:

Minispectrometer name	Spectral range, nm	Optical resolution, nm
HR2000 UV	247 - 473	0.17
HR 2000 VIS	460 - 663	0.15
HR 2000 H alpha	630 - 680	0.04
HR 2000 IR	659 - 1084	0.23

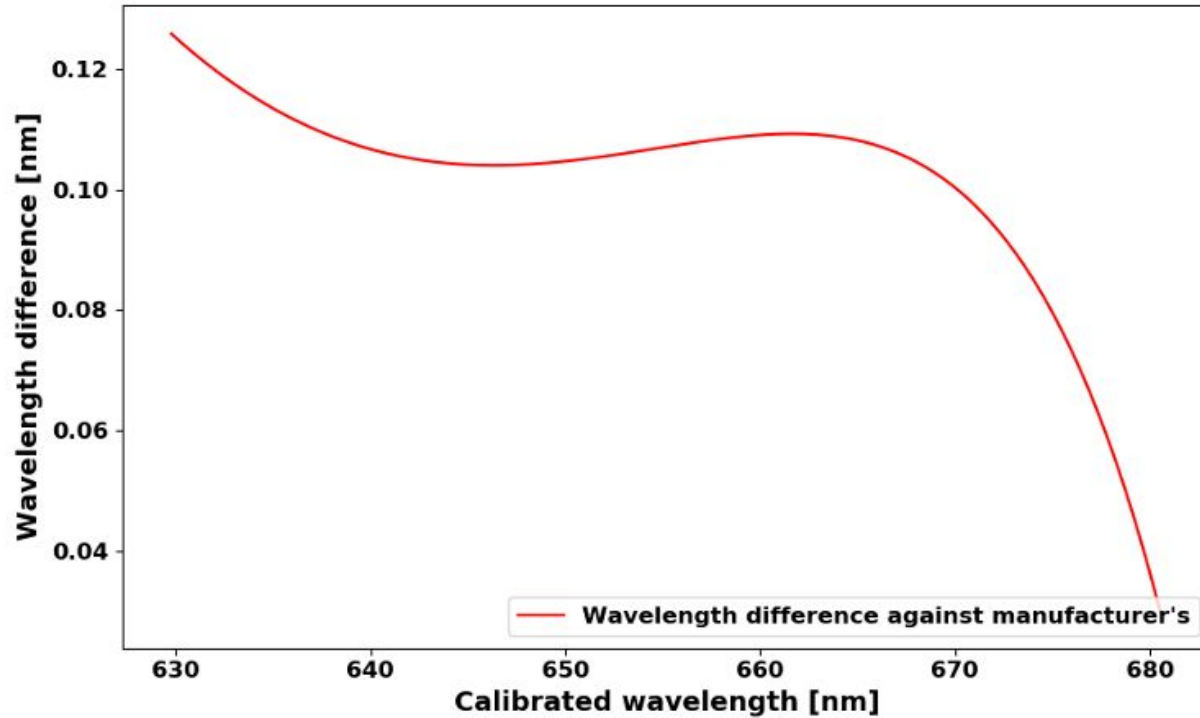
→ Calibration of Spectrometers!

allows precise wavelength assignment

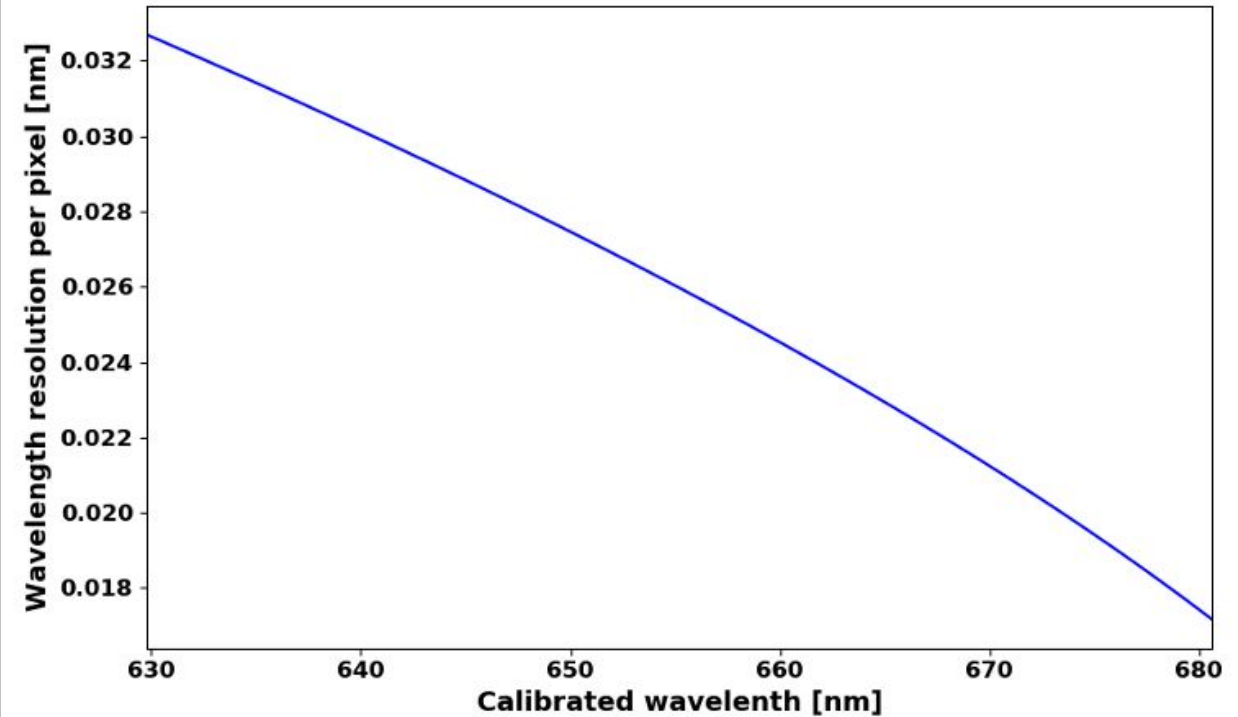
- **Hg-2 (Hg+Ar) spectral calibration source:**
Usable for UV, IR and VIS
- **Ne-2 spectral calibration source:**
Usable for VIS and Halpha



Halpha Spectrometer with Ne Source

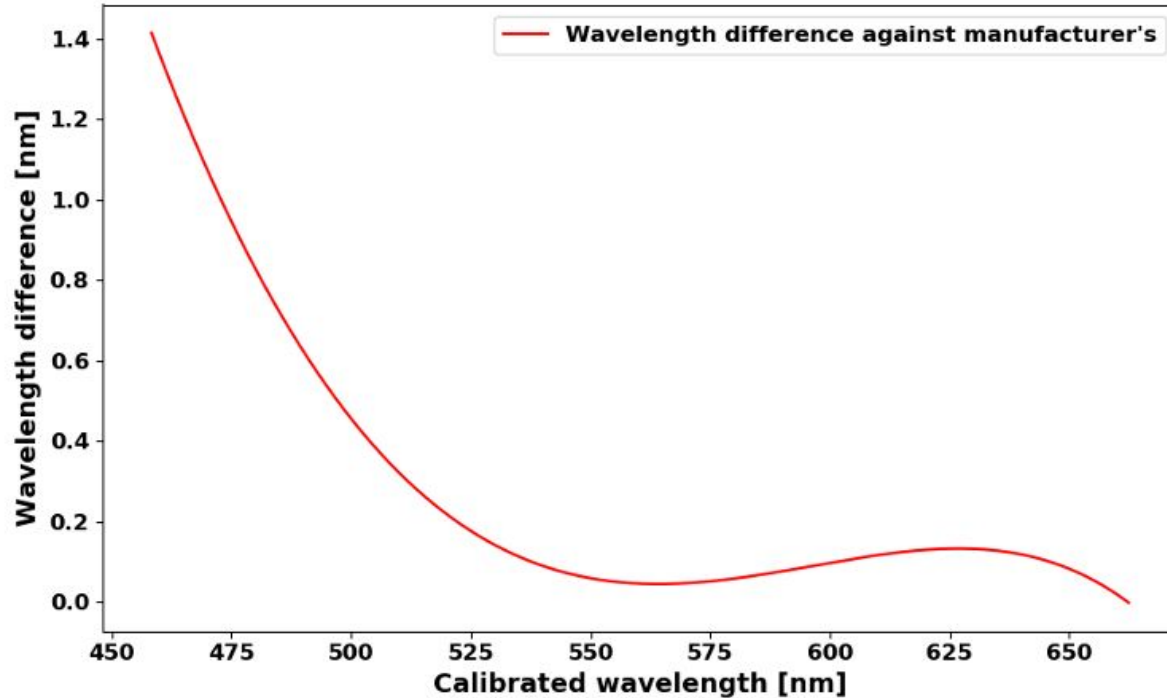


Halpha Spectrometer with Ne Source

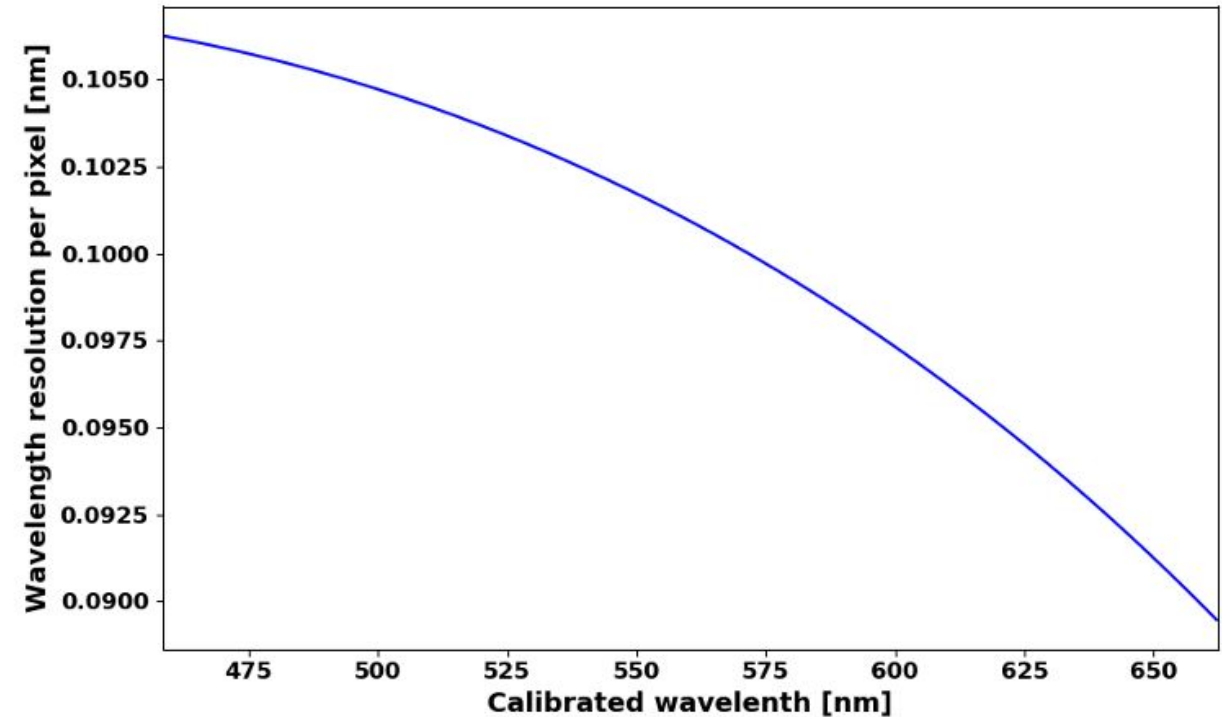


- Wavelength difference $\Delta\lambda \sim 0.1\text{nm}$ against the manufacturer's calibration
- Resolution per pixel $\sim 0.026\text{ nm/pixel} \rightarrow \sim 4\text{ pixels systematic error} \rightarrow$ Use new spectral calibration!
- FWHM $\sim 0.062\text{ nm}$

VIS Spectrometer with Hg and Ne Sources

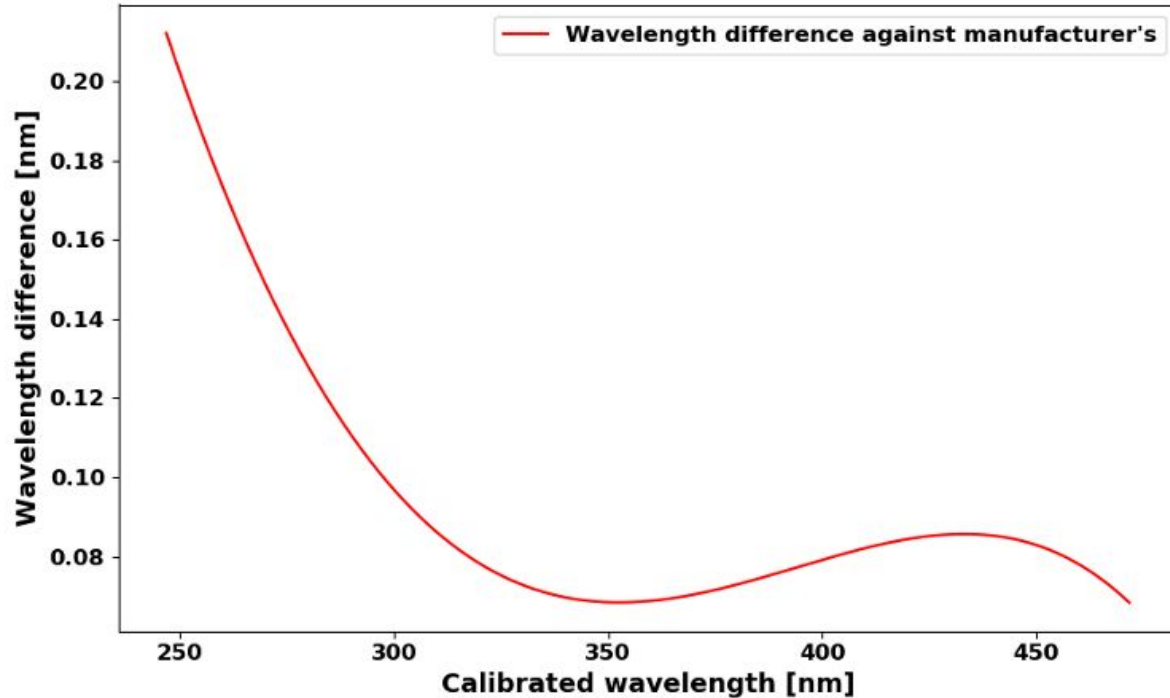


VIS Spectrometer with Hg and Ne Sources

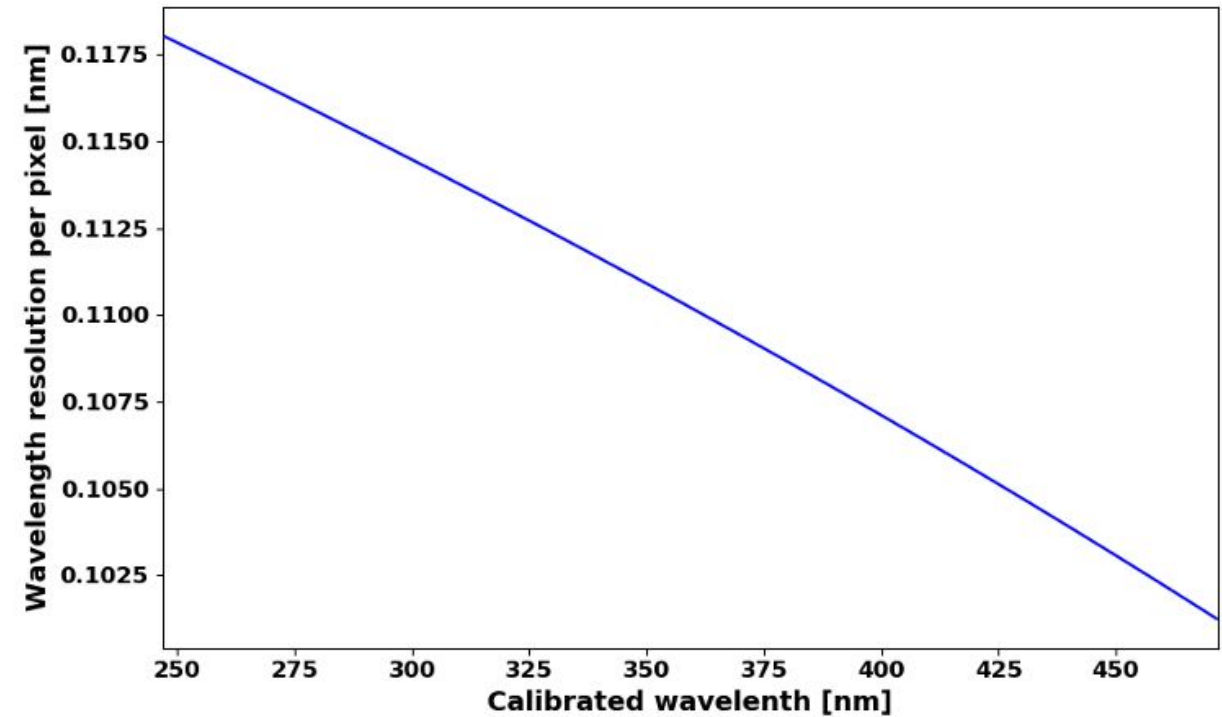


- **Wavelength difference negligible** (at 545-660 nm) against the manufacturer's calibration
 - Resolution per pixel ~ 0.1 nm/pixel
 - FWHM ~ 0.23 nm
- **Source is non ideal for this spectrometer**
Use manufacturer's calibration

UV Spectrometer with Hg Source

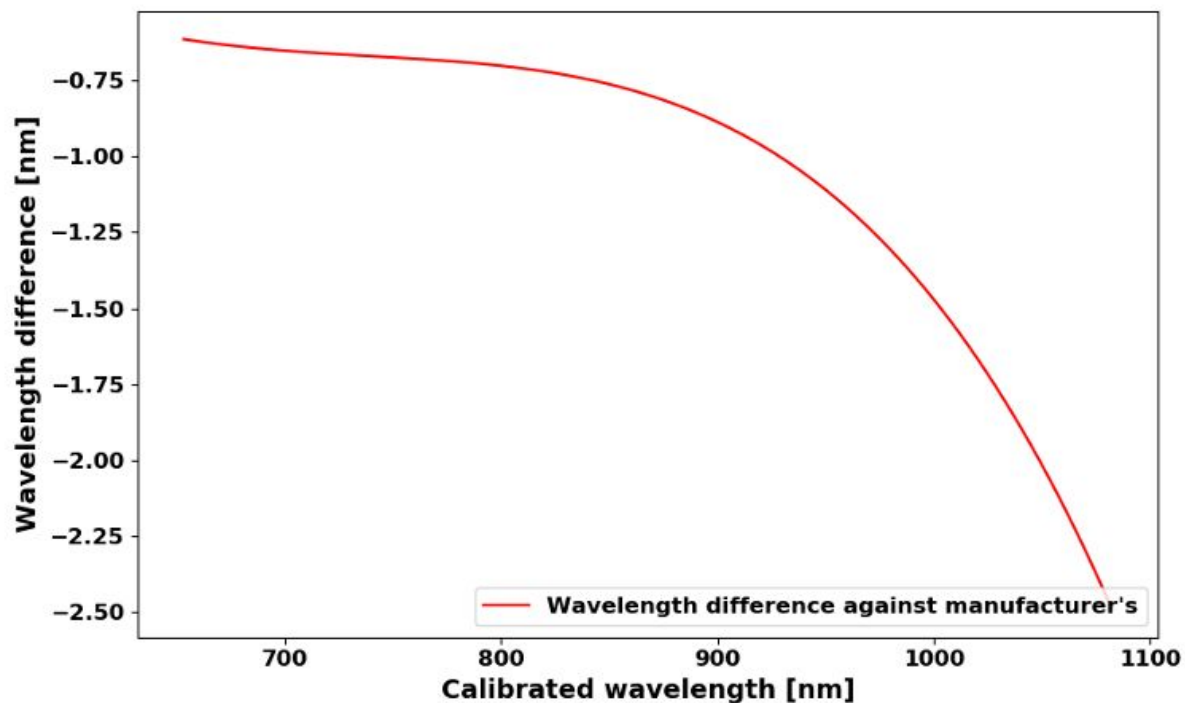


UV Spectrometer with Hg Source

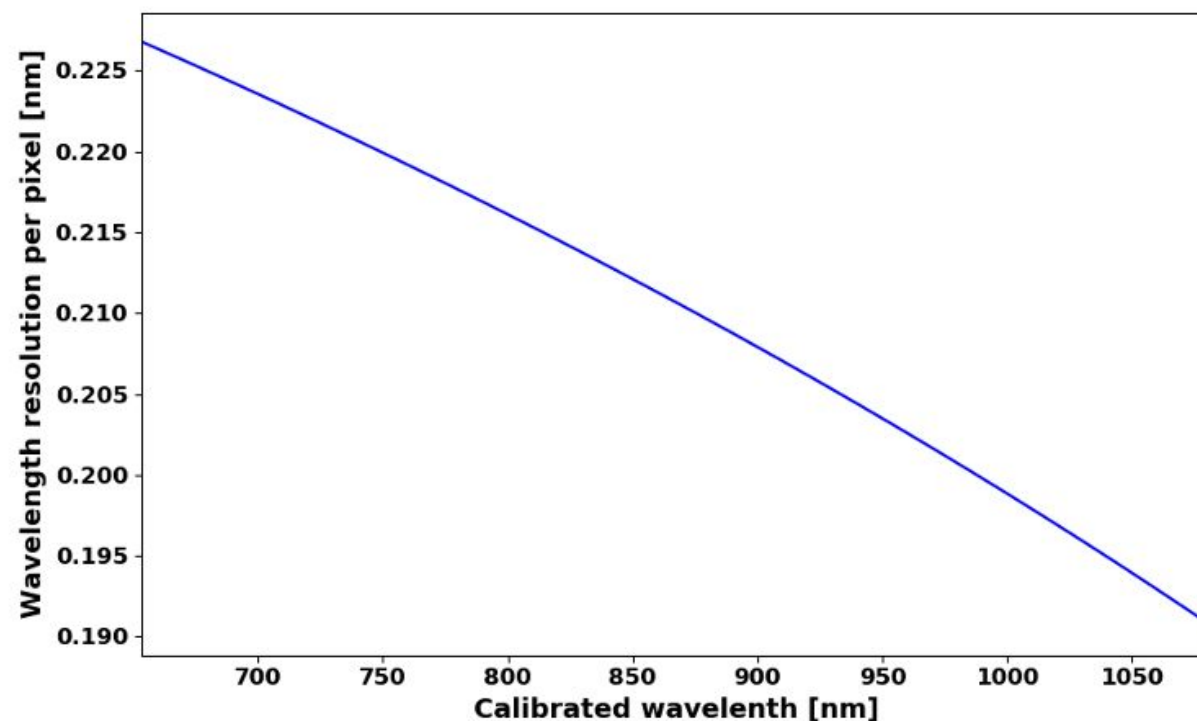


- Wavelength difference $\Delta\lambda \sim 0.07\text{nm}$ (at 296-436 nm) against the manufacturer's calibration
 - Resolution per pixel $\sim 0.1 \text{ nm/pixel}$
 - FWHM $\sim 0.23 \text{ nm}$
- Use manufacturer's calibration!

IR Spectrometer with Ar Source

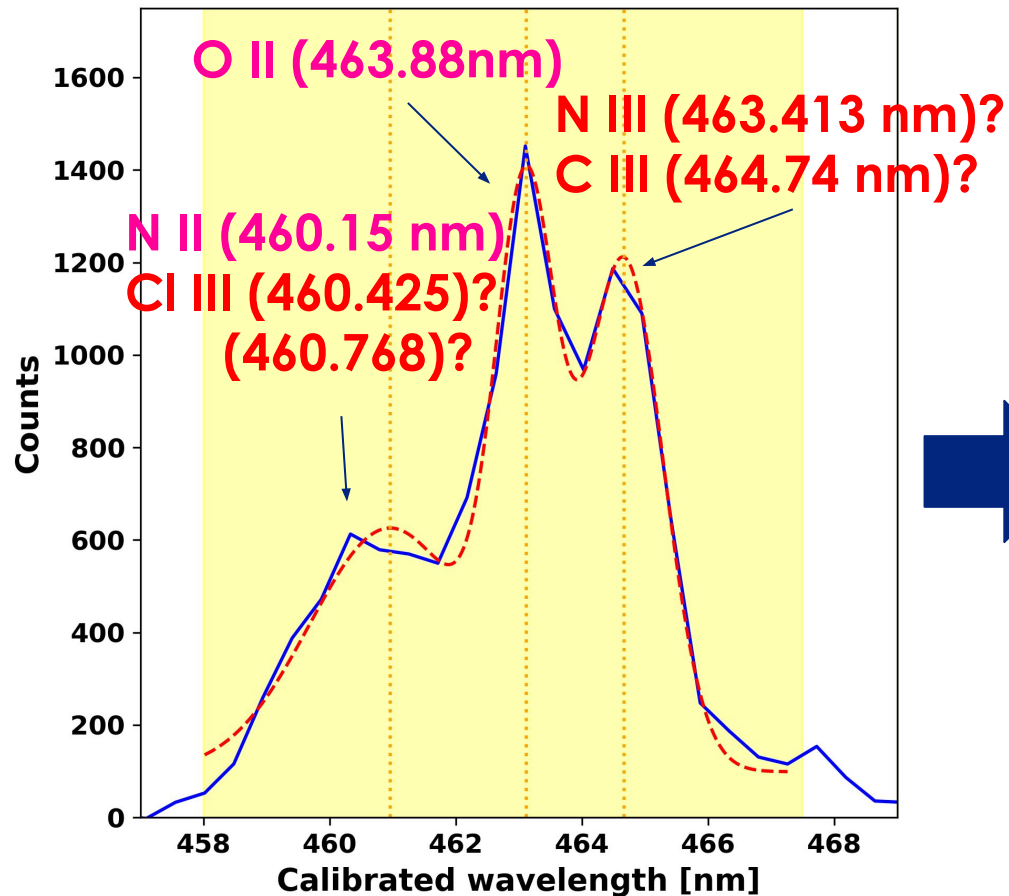


IR Spectrometer with Ar Source

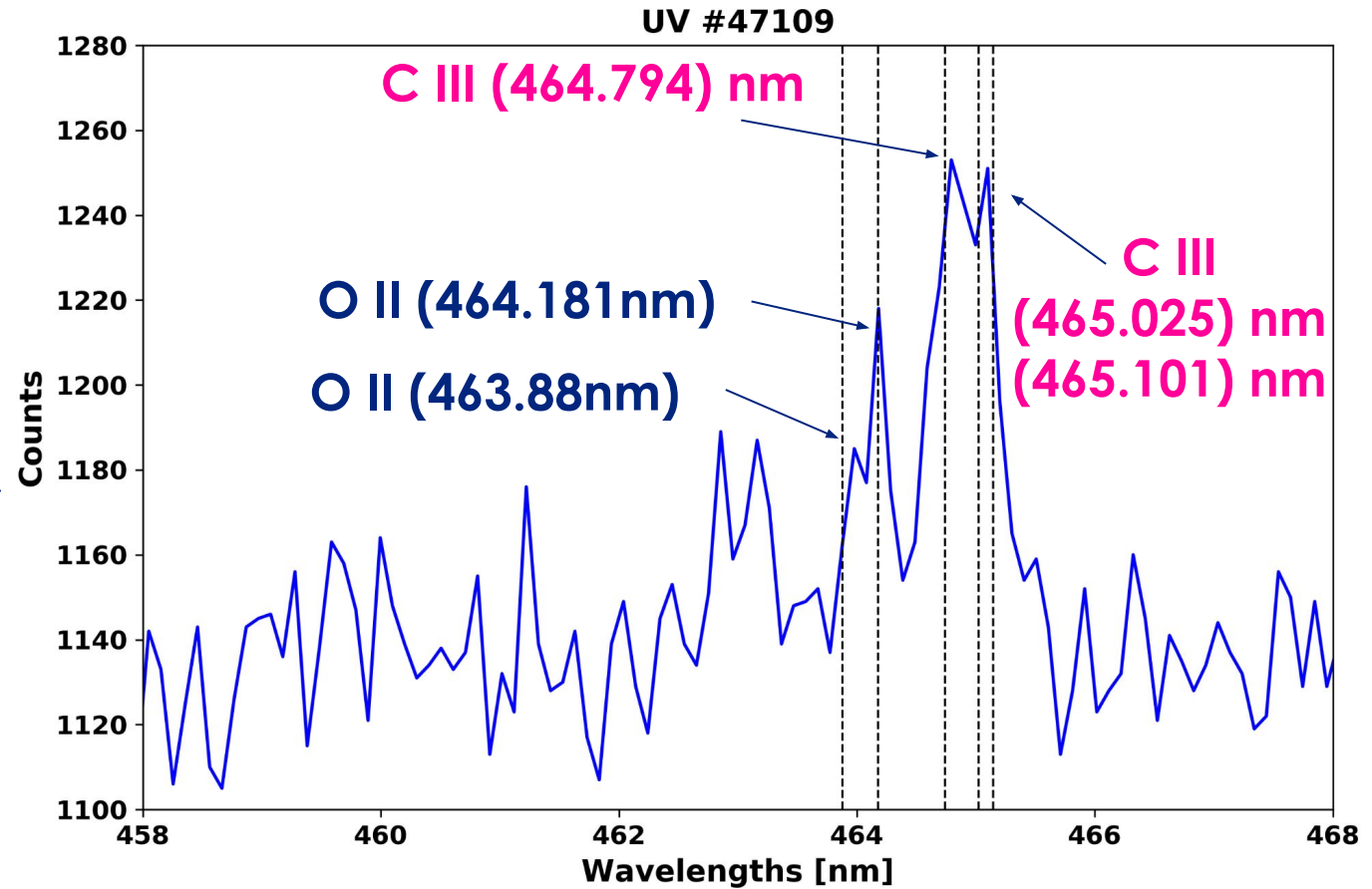


- Wavelength difference $\Delta\lambda \sim 0.75$ nm (at 696-966 nm) against the manufacturer's calibration
- Resolution per pixel ~ 0.1 nm/pixel $\rightarrow \sim 4$ pixels systematic error $\rightarrow$ Use new calibration up to 1000!
- FWHM ~ 0.39 nm

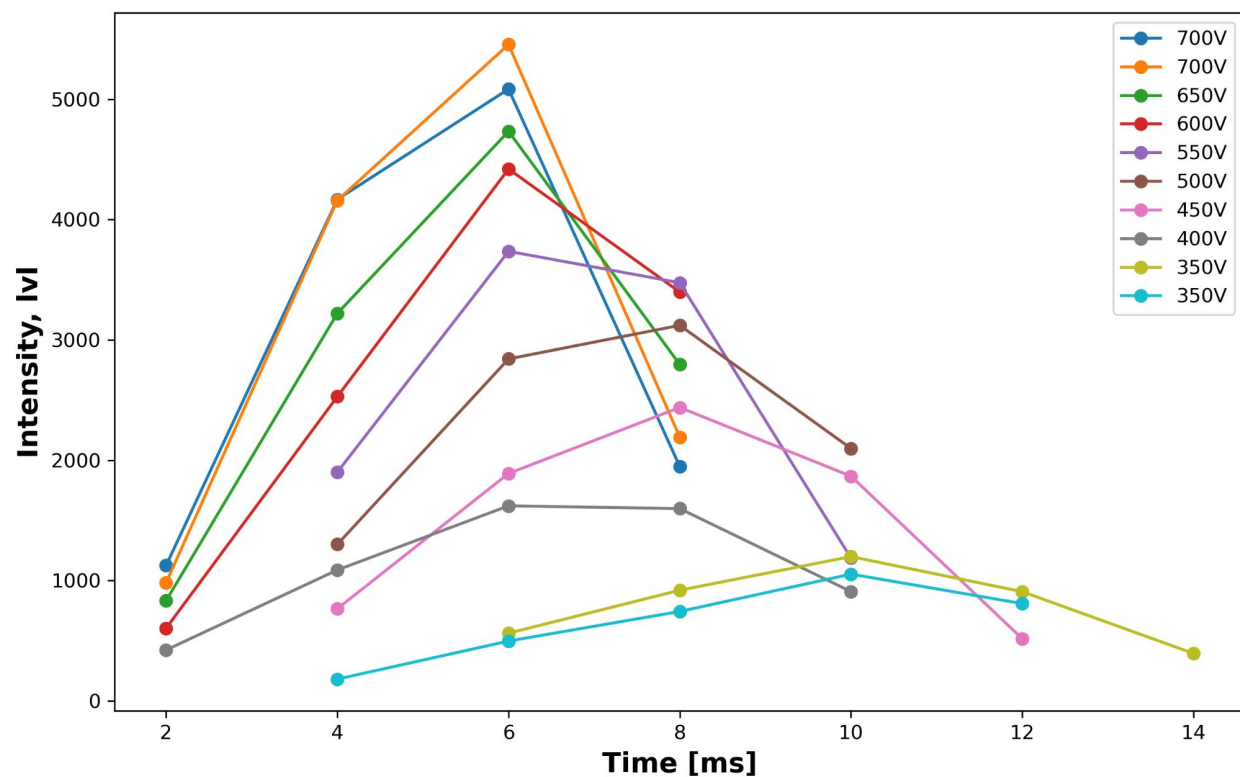
Previously (IRVISUV)



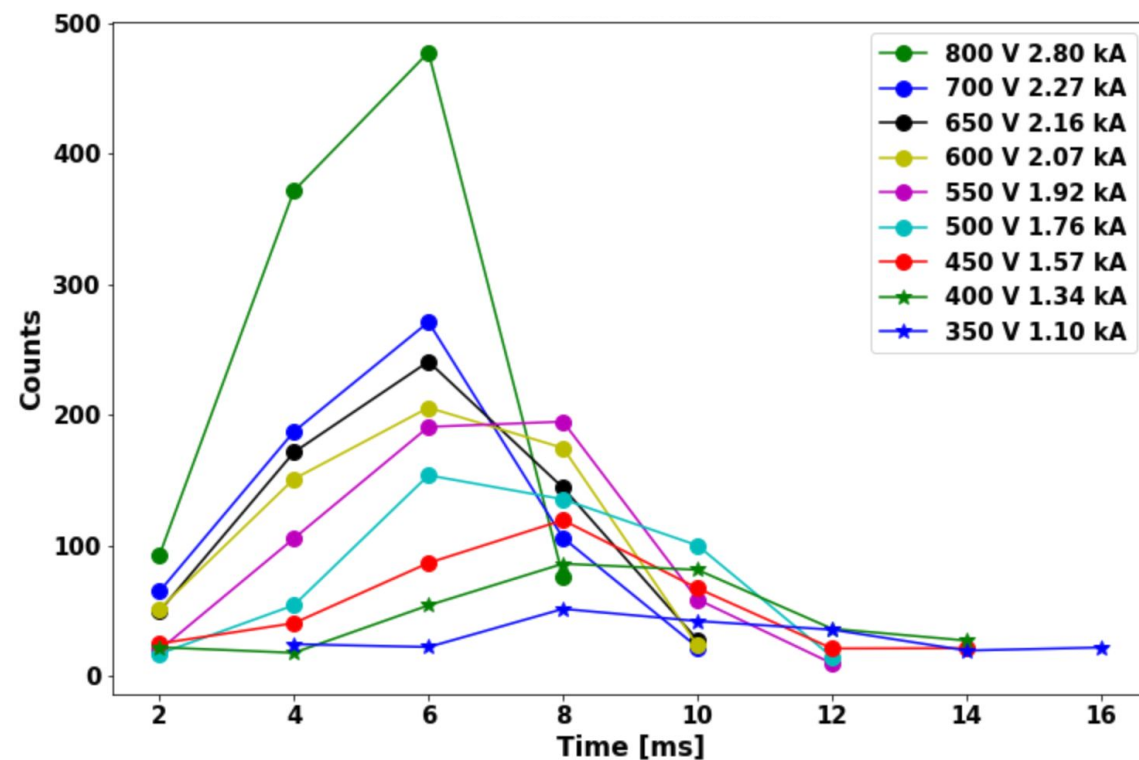
C III triplet FOUND!



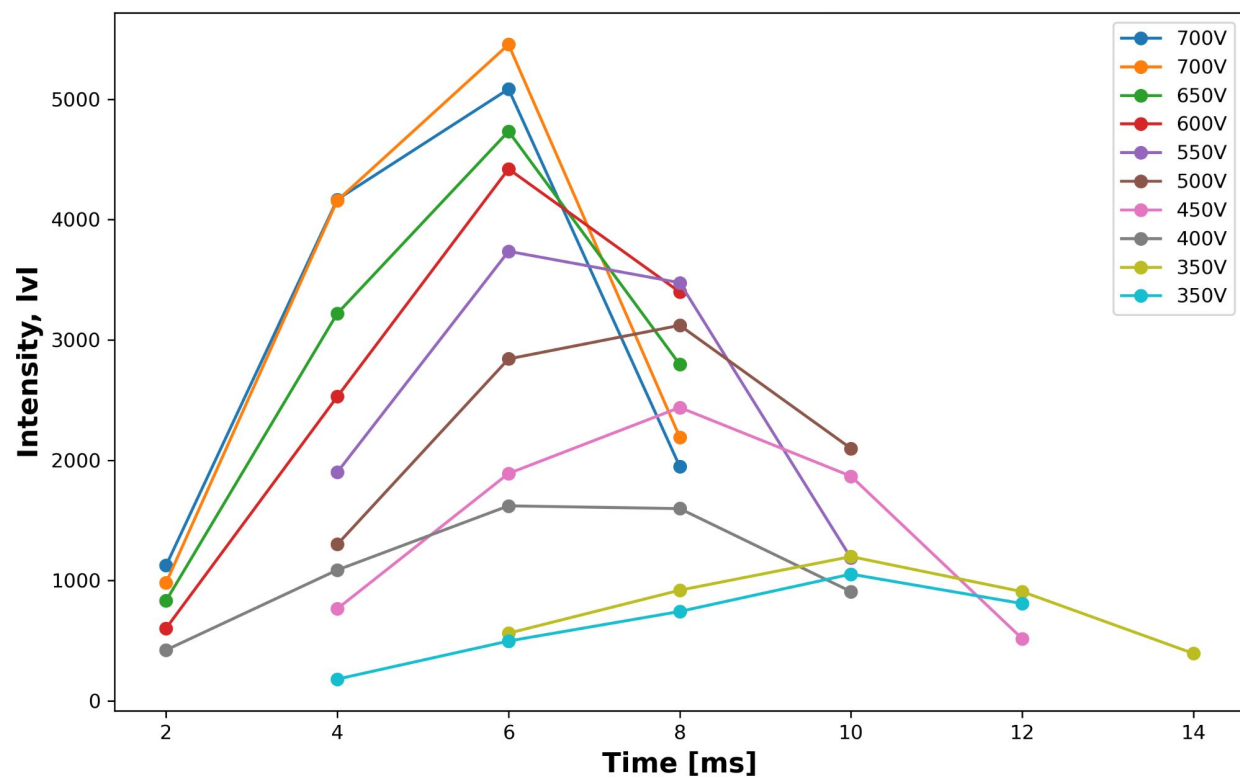
Big Peak Time Evolution (IRVISUV)



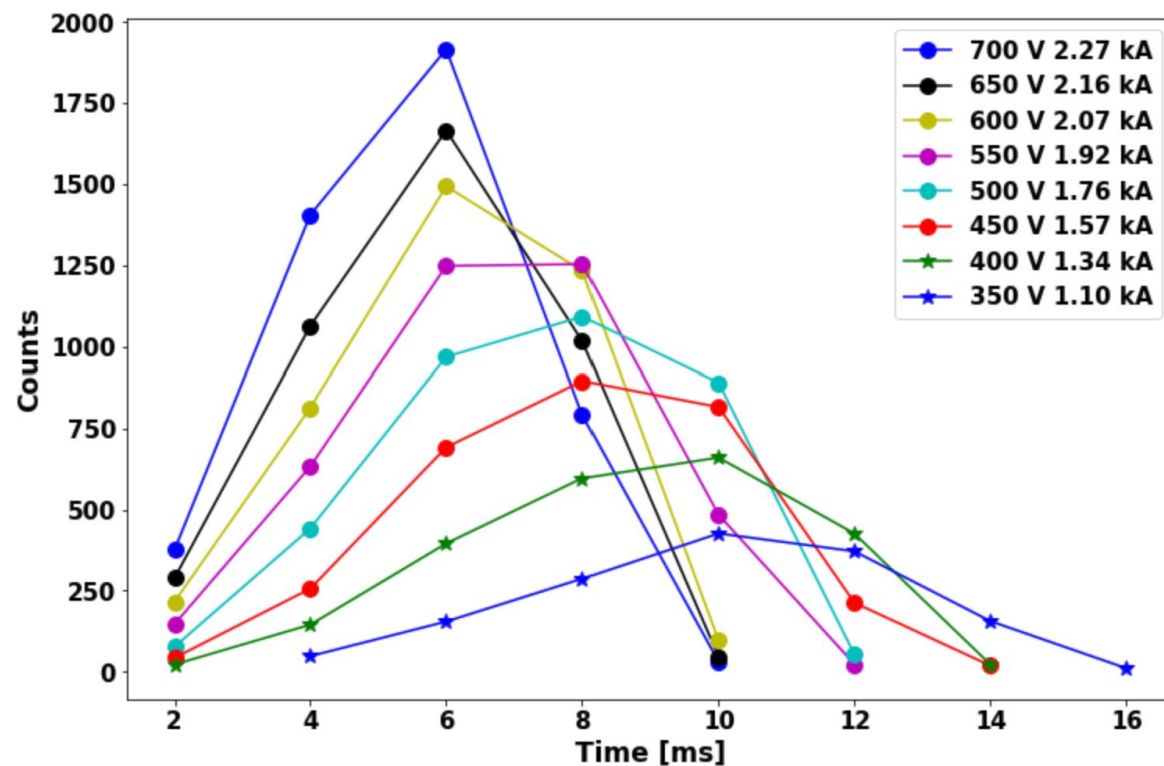
O II Time Evolution (IRVISUV)



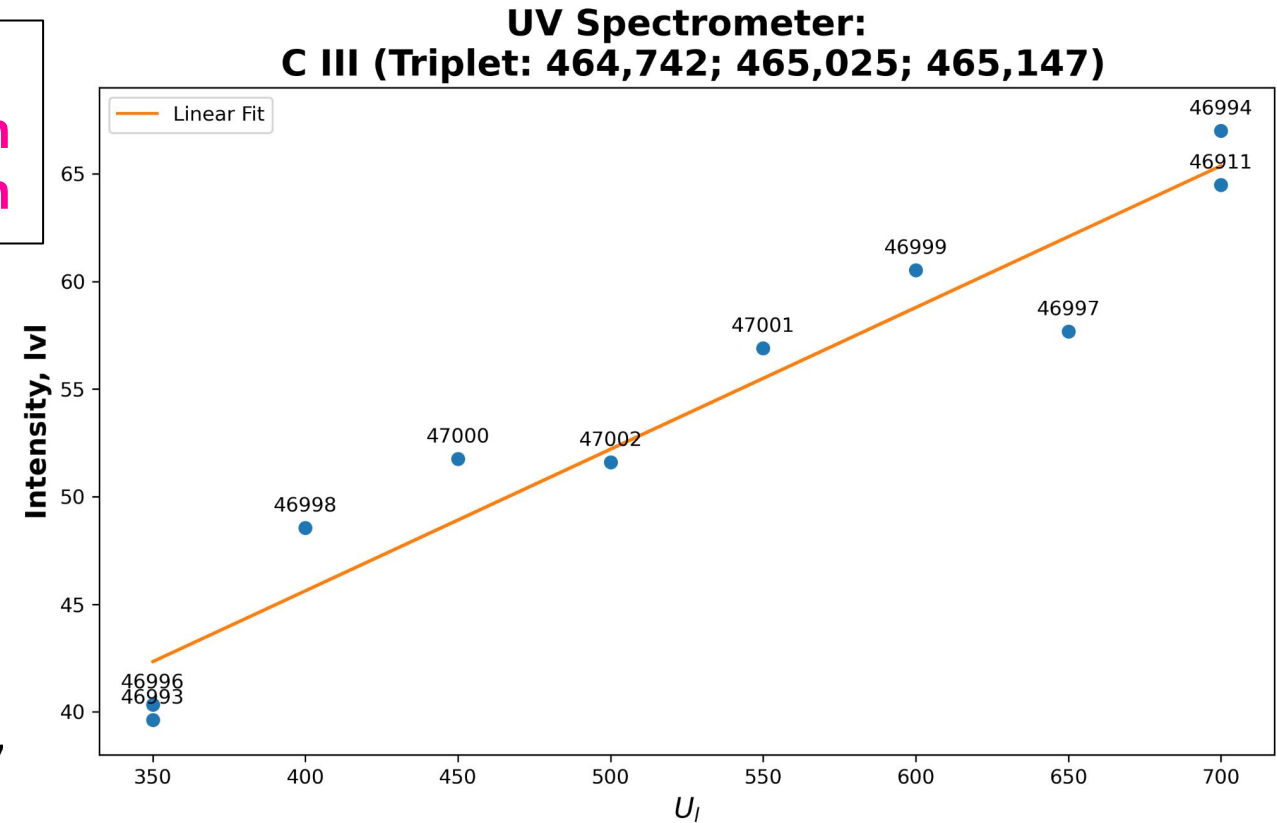
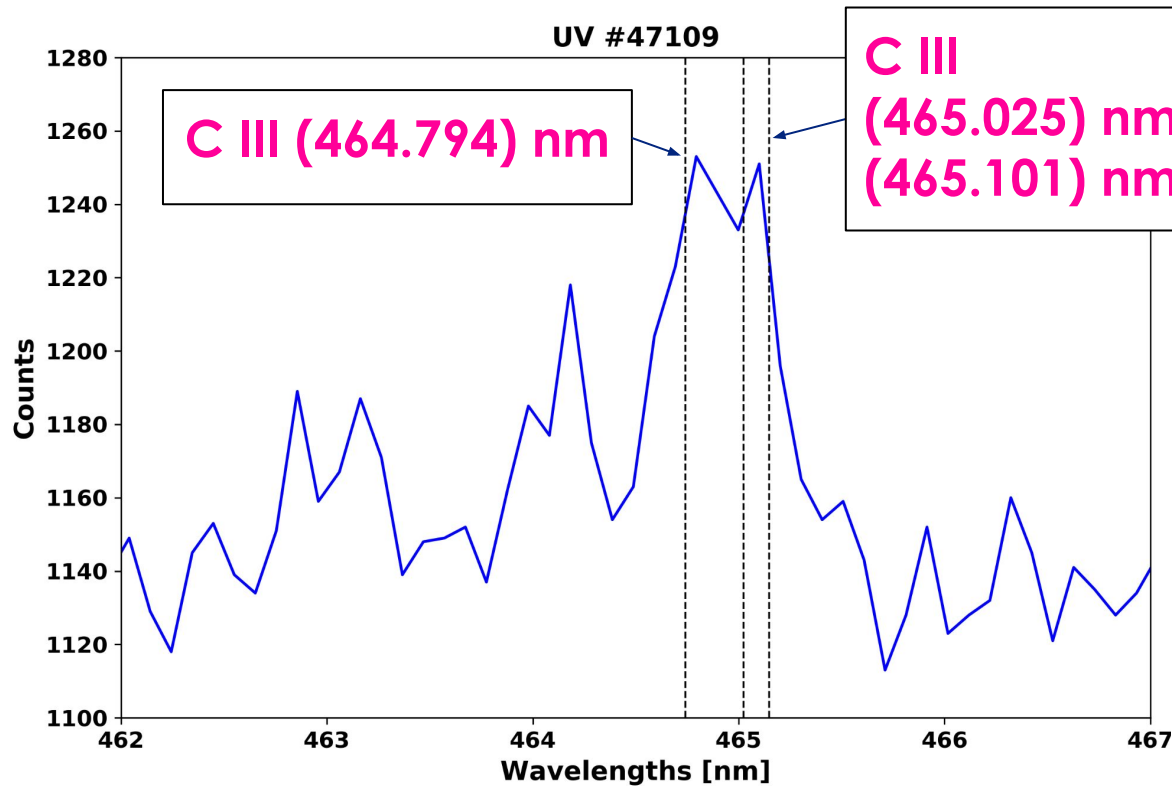
Big Peak Time Evolution (IRVISUV)



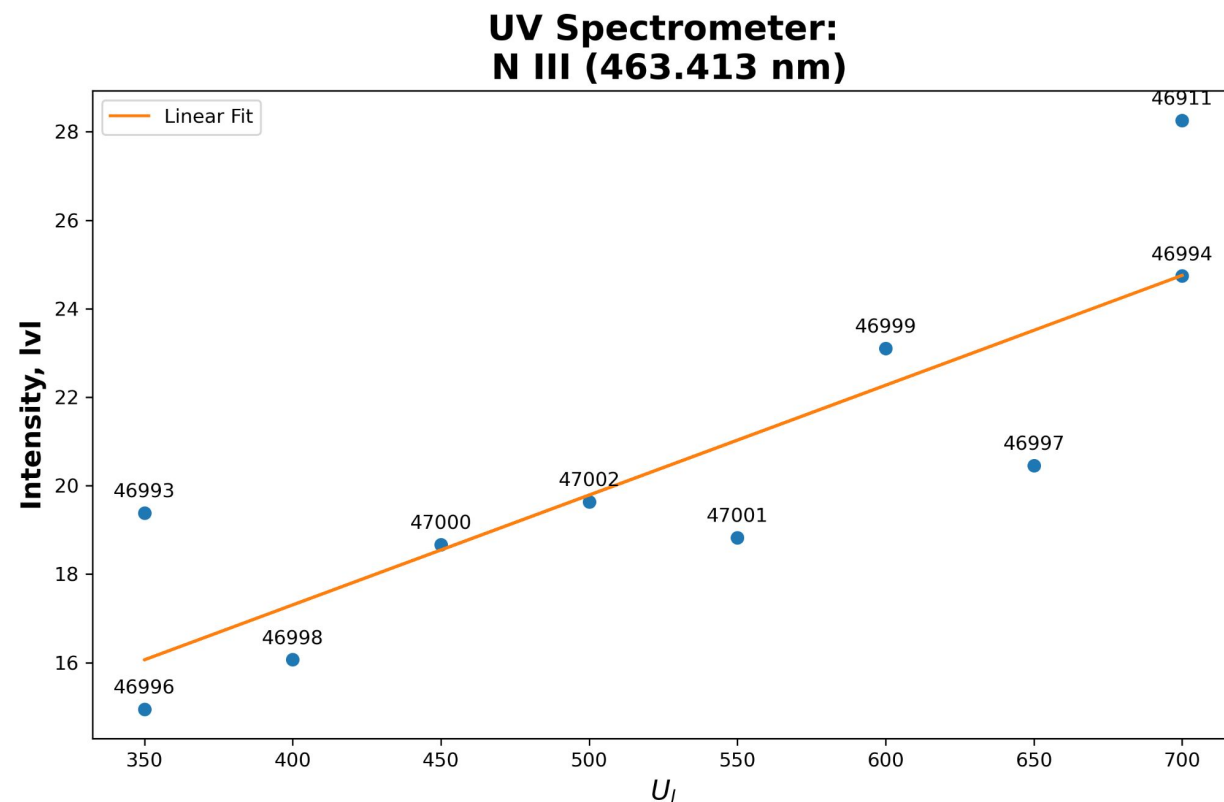
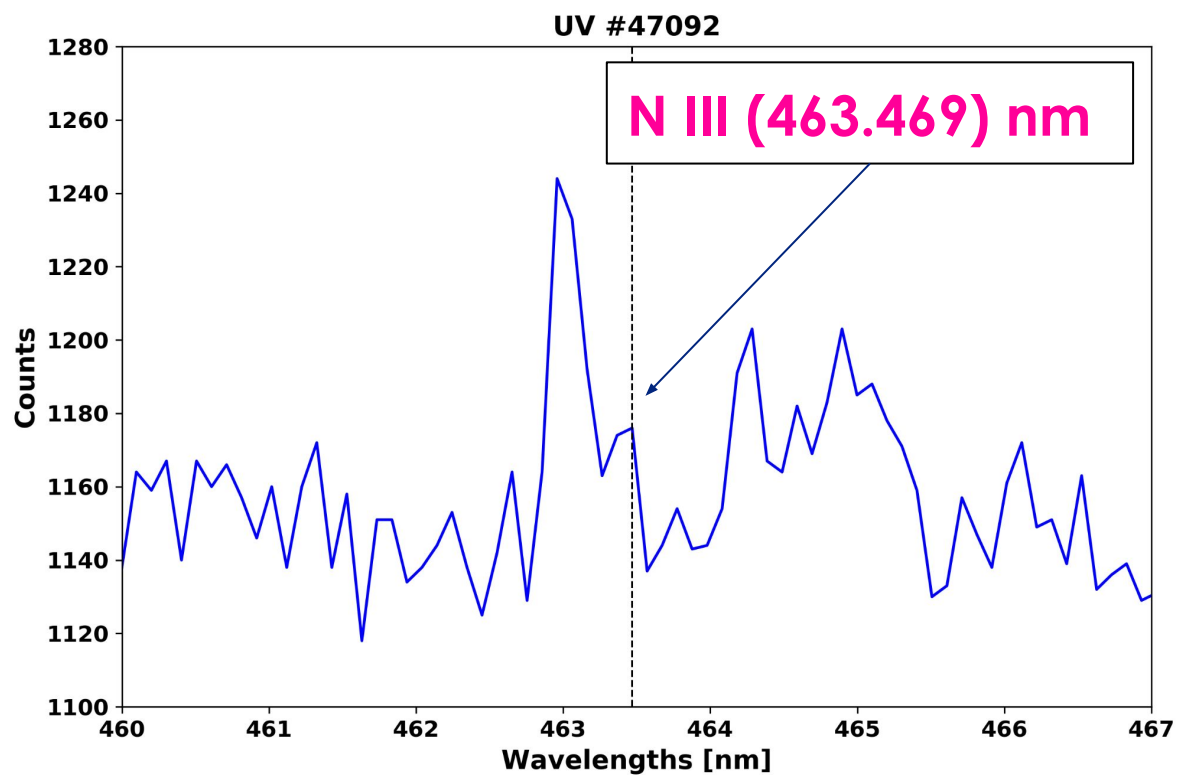
N II Time Evolution (IRVISUV)



C III triplet: 464.742, 465.025, 465.147 nm (UV: $\Delta\lambda=0.17$ nm)



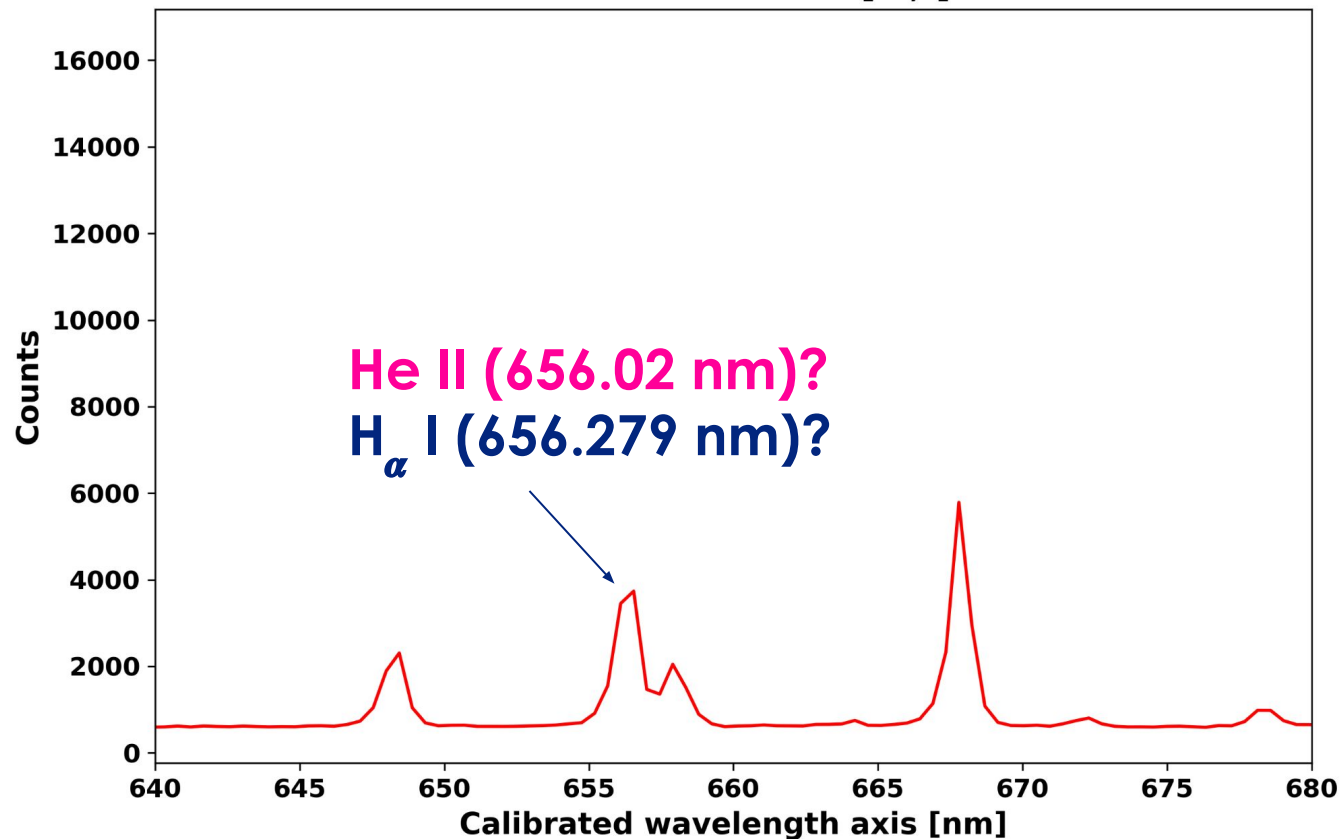
N III doublet?: 463.413 nm



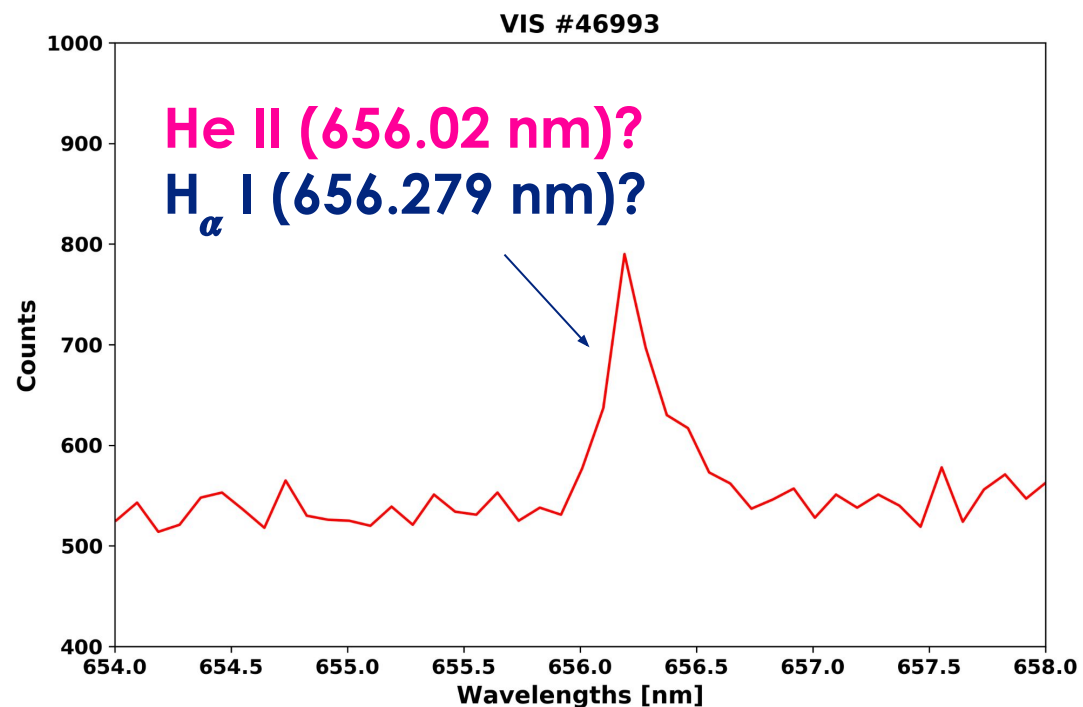
He II?

Previously:

IRVISUV #46991 ms [f3,r]

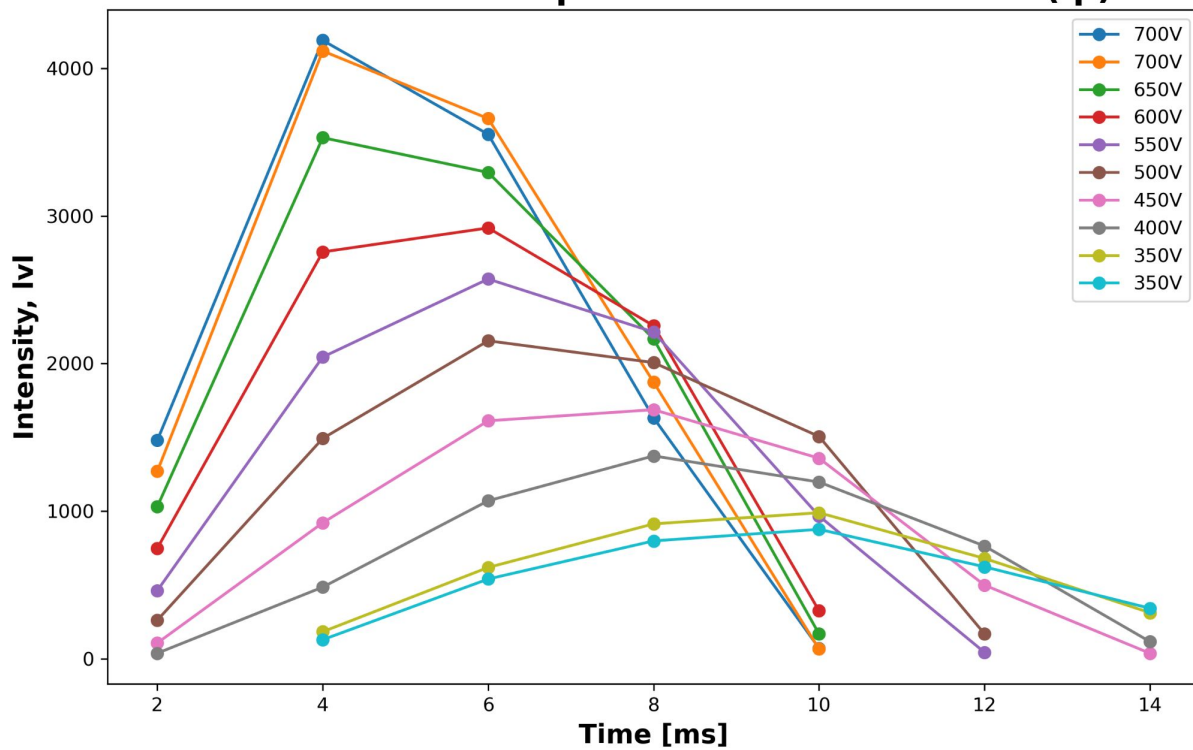


VIS: $\Delta\lambda=0.15$ nm

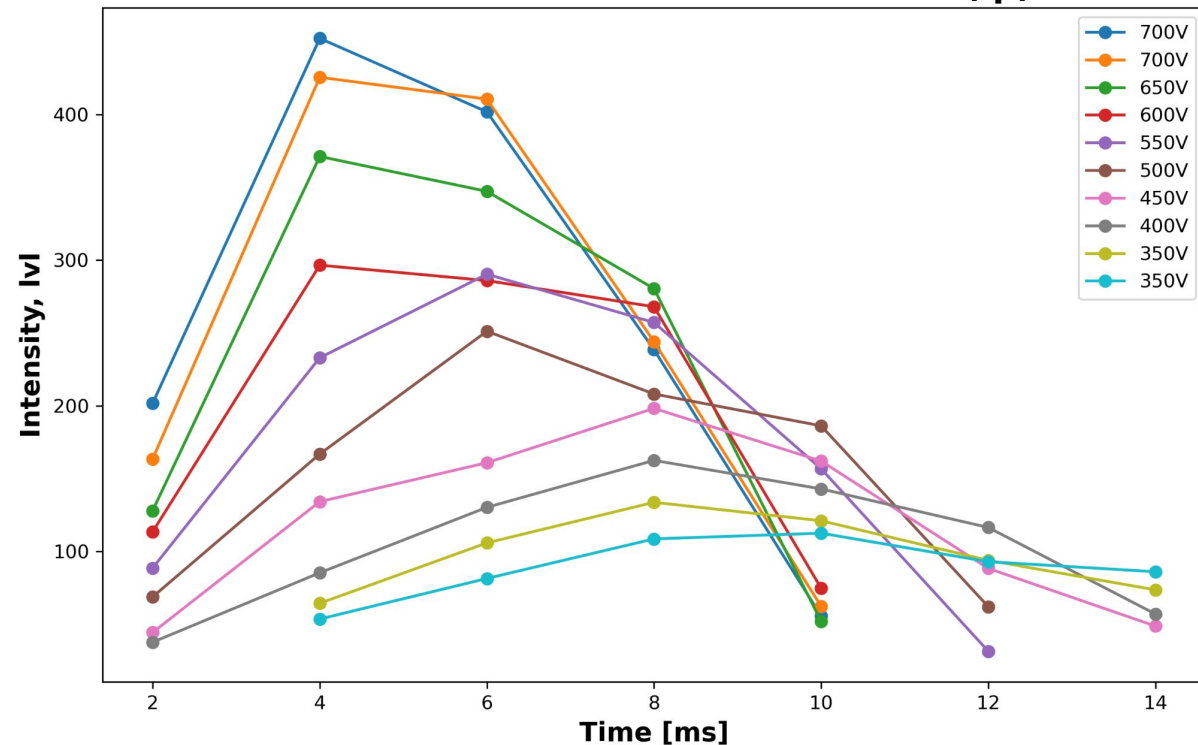


Comparison with H_{β} (486.135 nm)

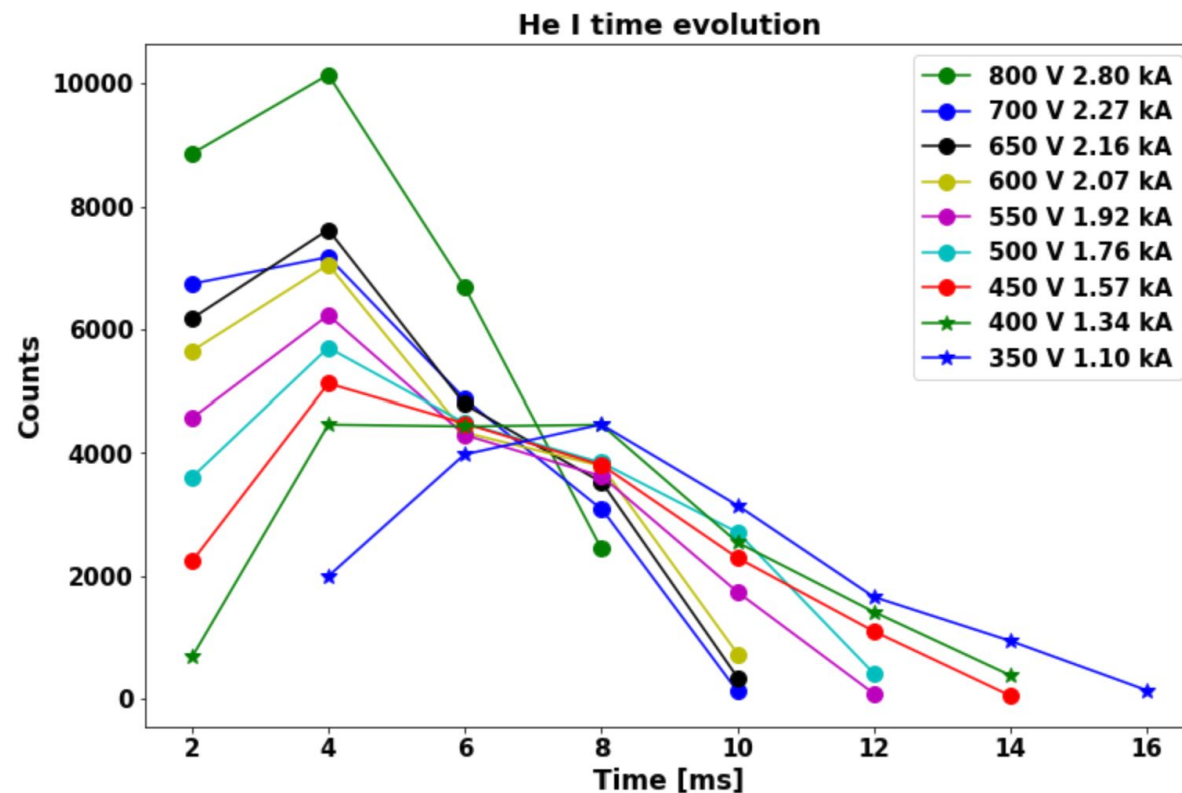
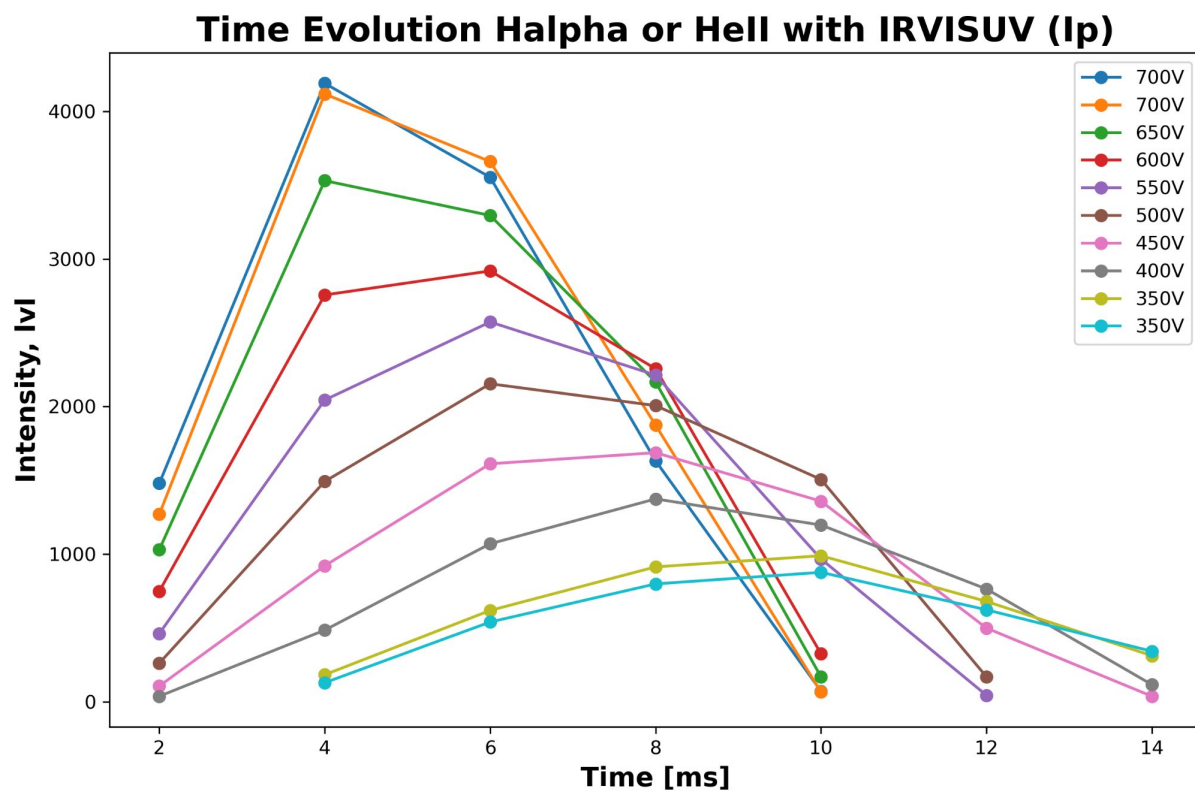
Time Evolution Halpha or HeII with IRVISUV (Ip)



Time Evolution Hbeta with IRVISUV (Ip)

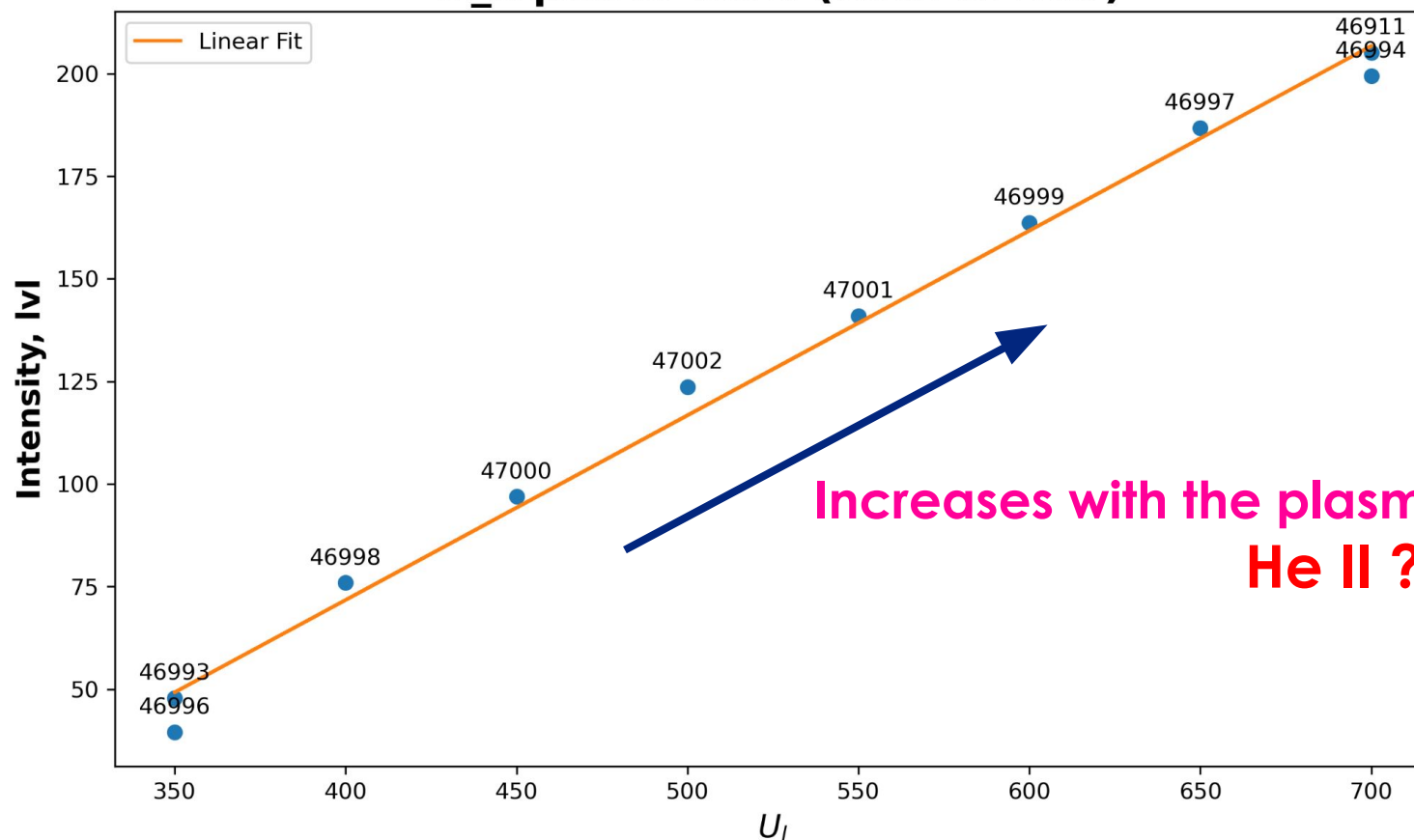


Comparison with He I (667.8 nm)



He II (656.02 nm) FOUND!

VIS Spectrometer:
H_α or He II (656.181 nm)

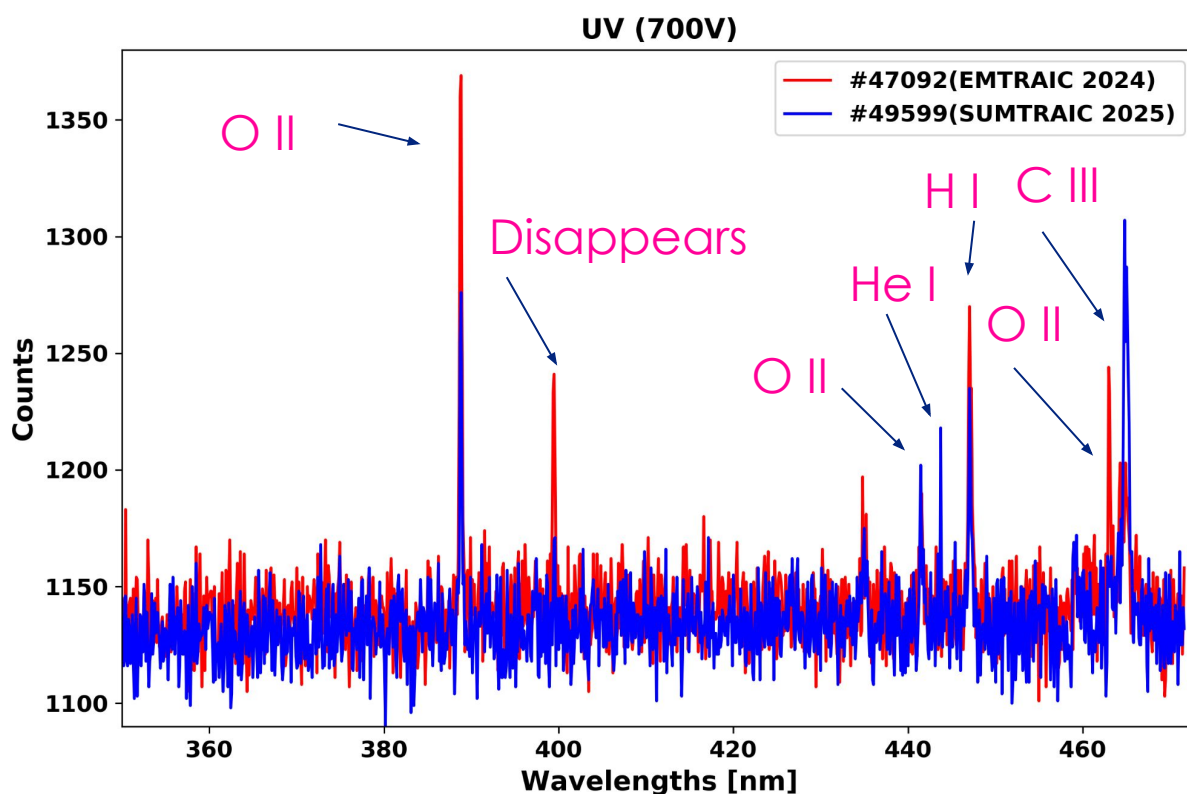


Increases with the plasma current → ion
He II ? 

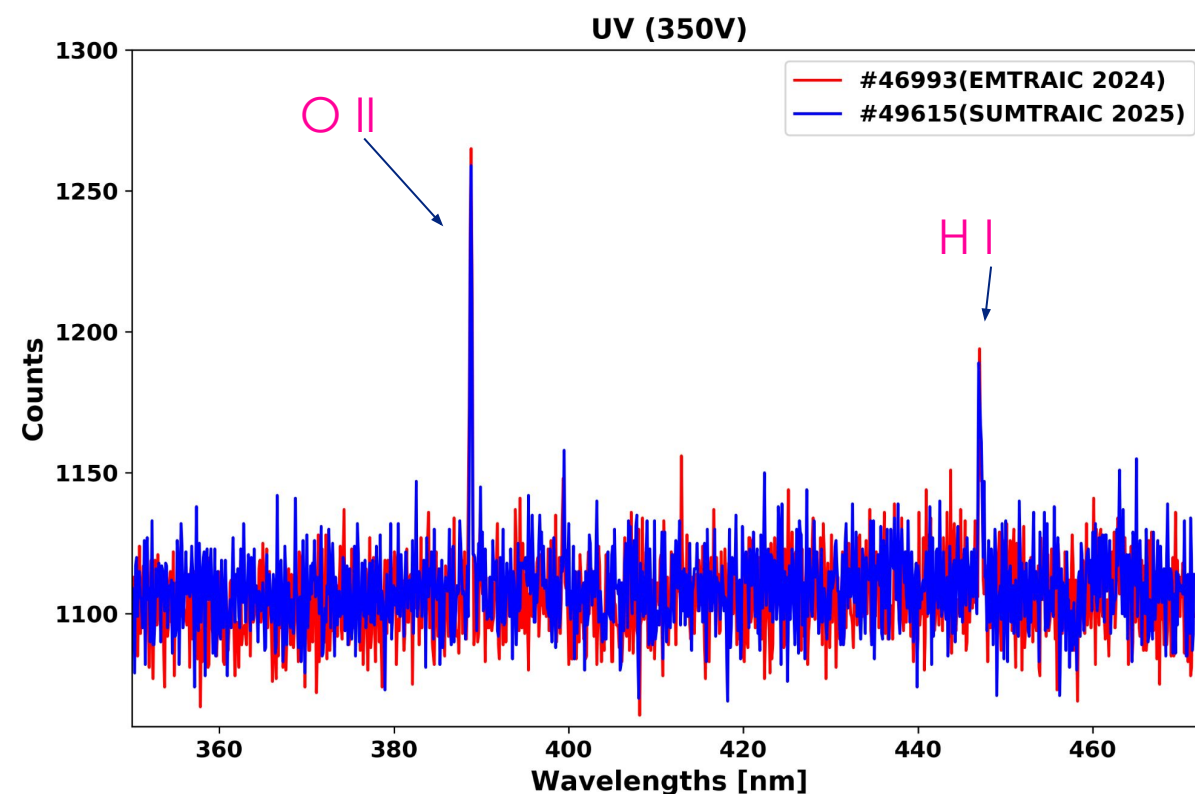
EMTRAIC (CABLE) VS SUMTRAIC (NO CABLE)



Transition:

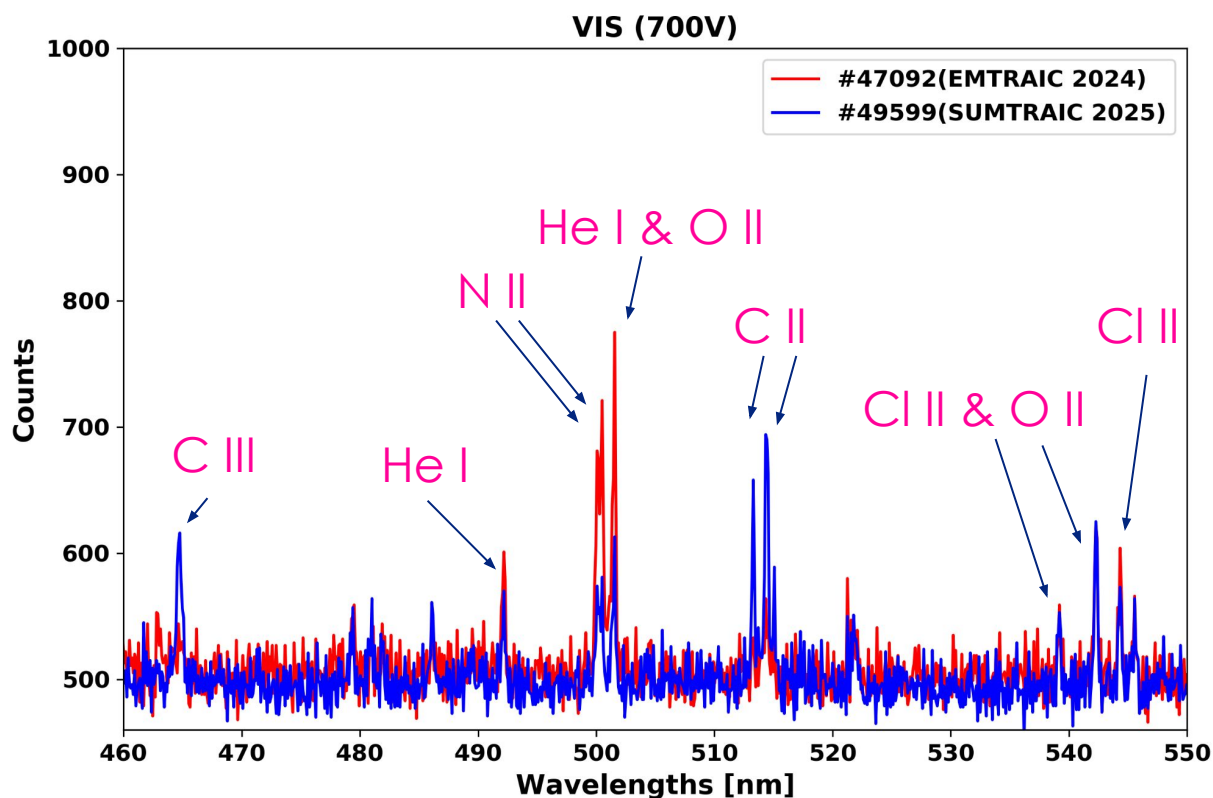


No Transition:

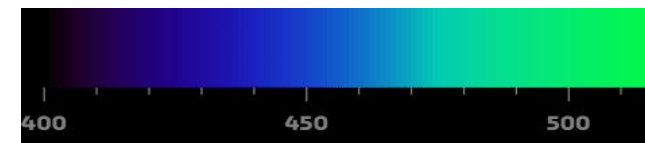
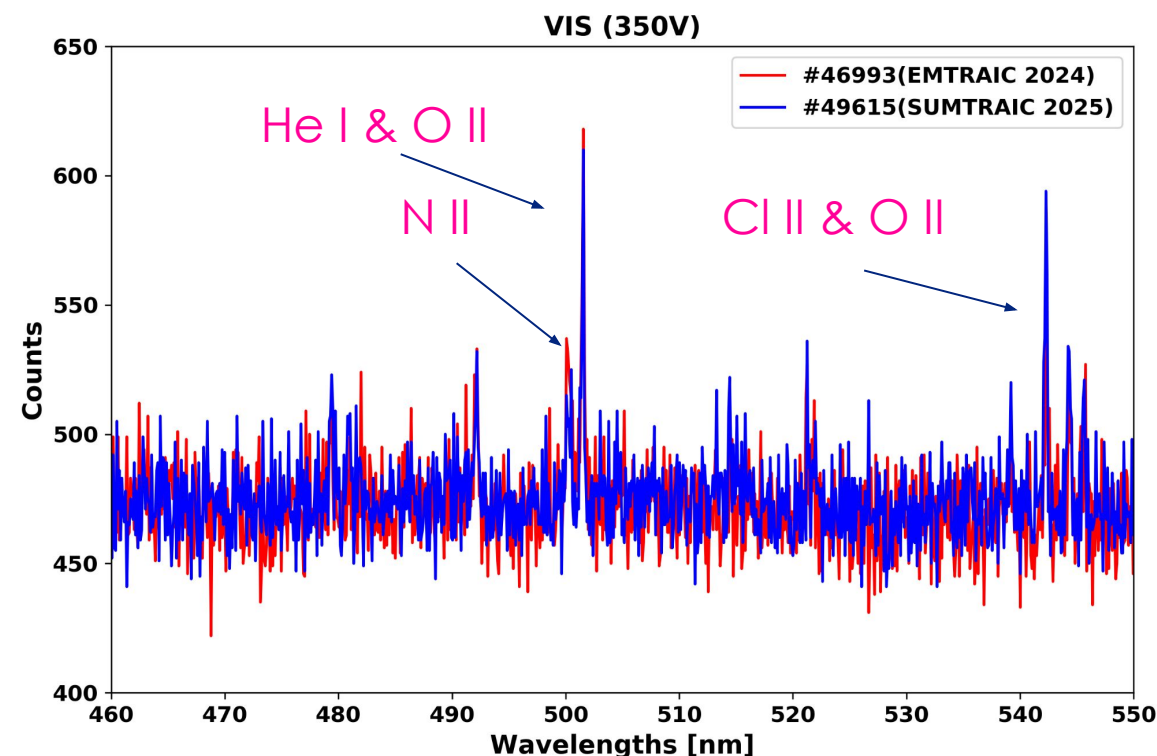


EMTRAIC (CABLE) VS SUMTRAIC (NO CABLE)

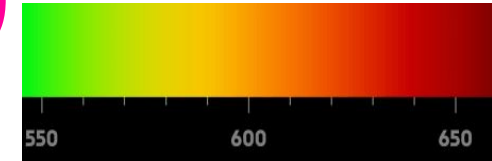
Transition:



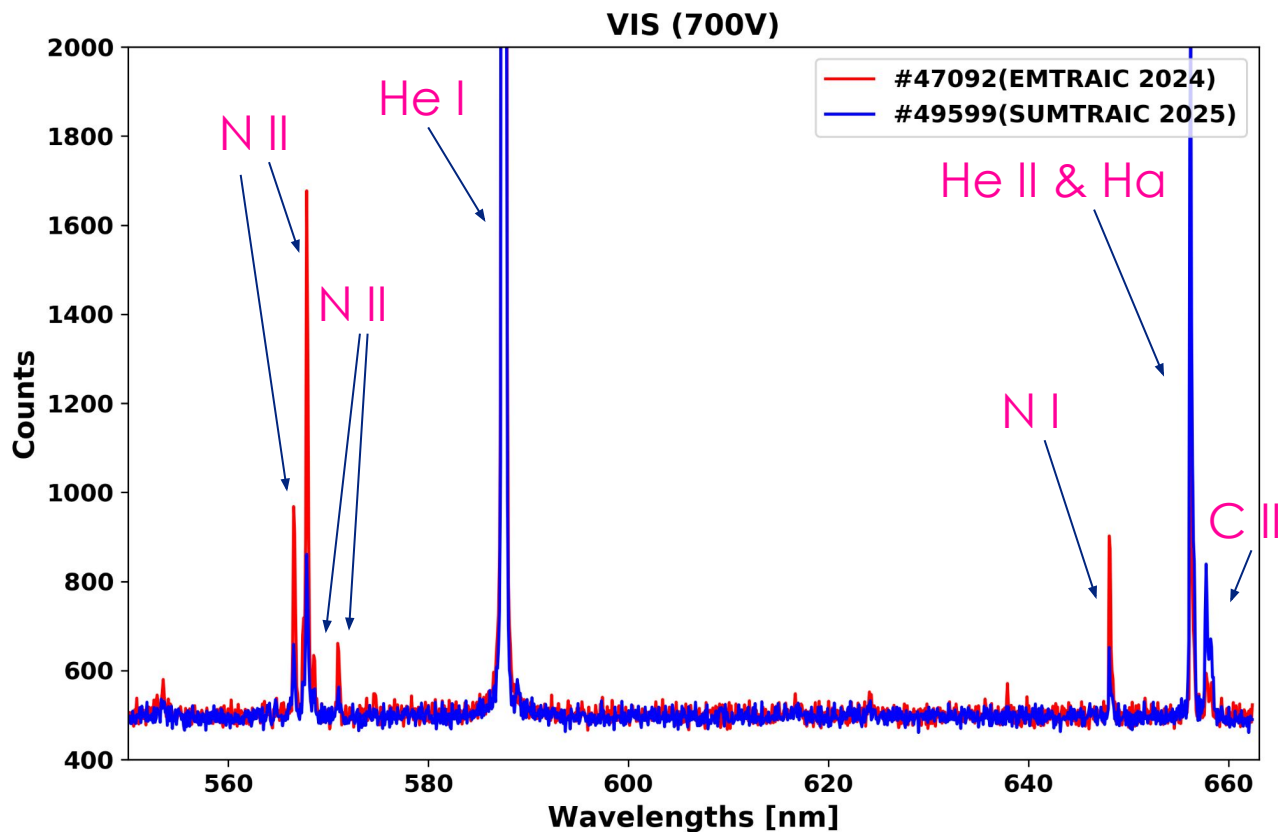
No Transition:



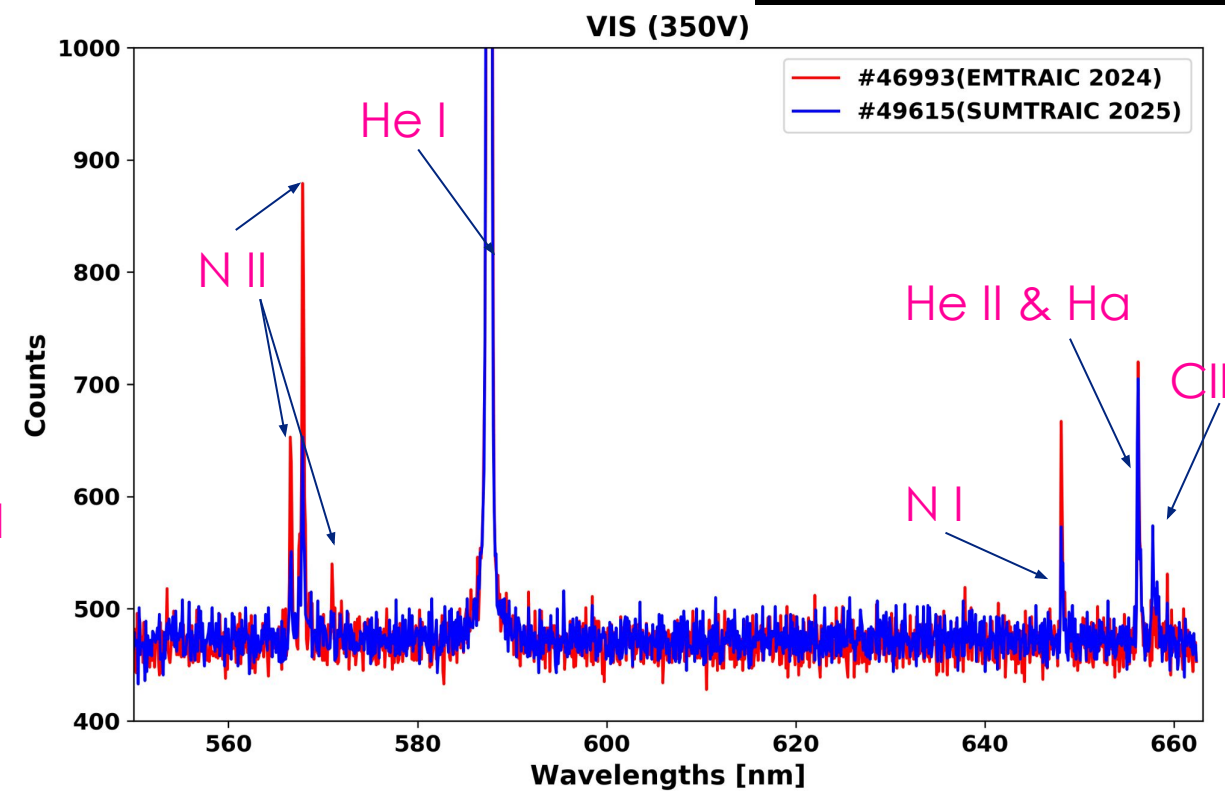
EMTRAIC (CABLE) VS SUMTRAIC (NO CABLE)



Transition:



No Transition:



Conclusion

- Found
- Not found?

Neutral (I)	Ionization energy [eV]	1st ionic state (II)	Ionization energy [eV]	2nd ionic state (III)
He	24.59	He ⁺		
H	13.60			
C	11.26	C ⁺	24.38	C ²⁺
O	13.62	O ⁺	35.12	O ²⁺
N	14.53	N ⁺	29.60	N ²⁺
Cl	12.97	Cl ⁺	23.8	Cl ²⁺

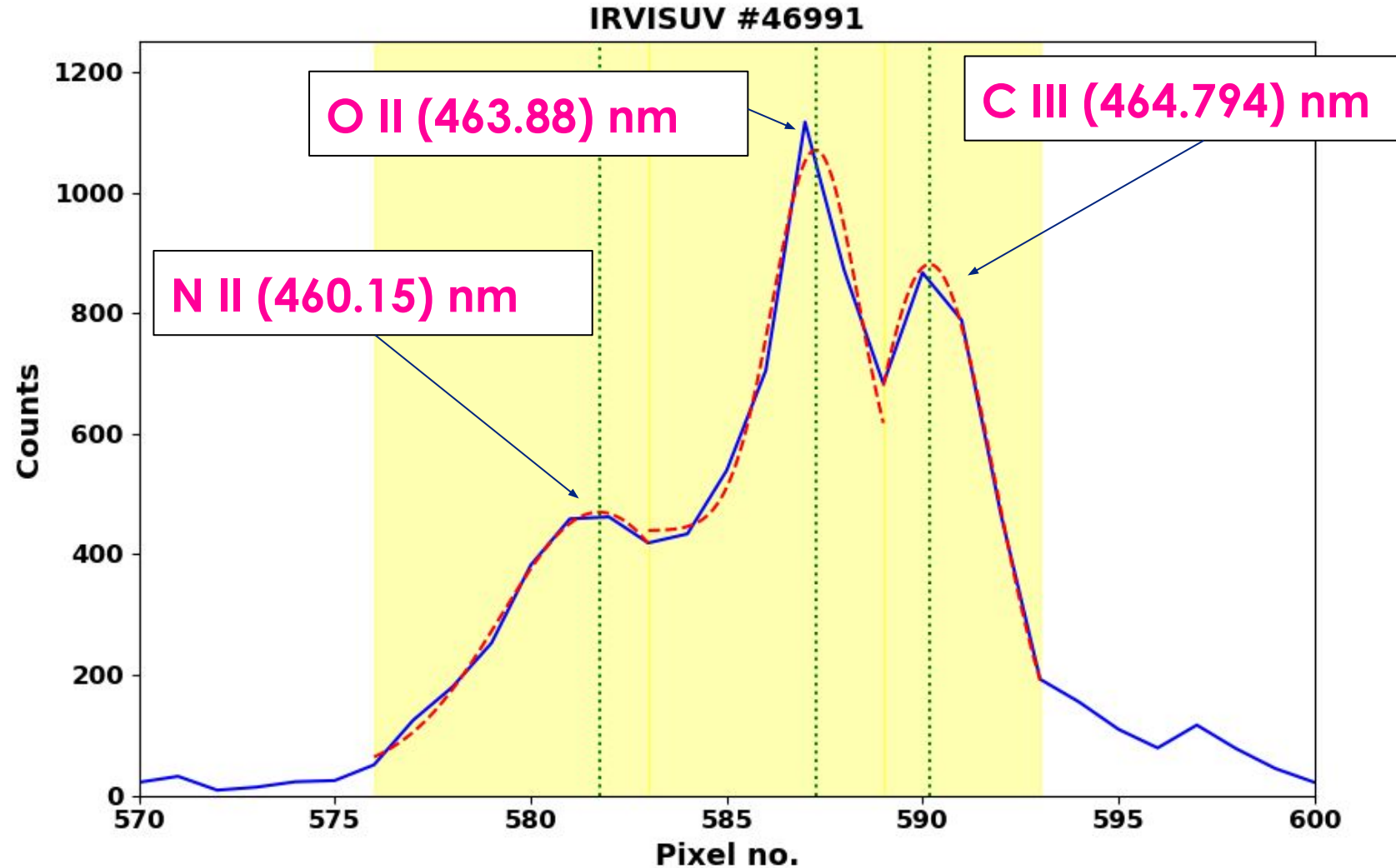
- Spectroscopy diagnostic system should be updated → not enough QE of high resolution spectrometers .
- Occasional presence of power cord → Impurity influx + Edge condition change → Repeat experiment
- Similar improvements in particle confinement were reported by Finken et al. (2007) in TEXTOR, where opening of flux tubes connected deeper rational surfaces to the wall led to edge cooling, core density peaking, and barrier formation.
- Our results suggest that even in a small device like GOLEM, analogous barrier-formation mechanisms may be at play, with impurities serving as tracers of changing edge topology.

References

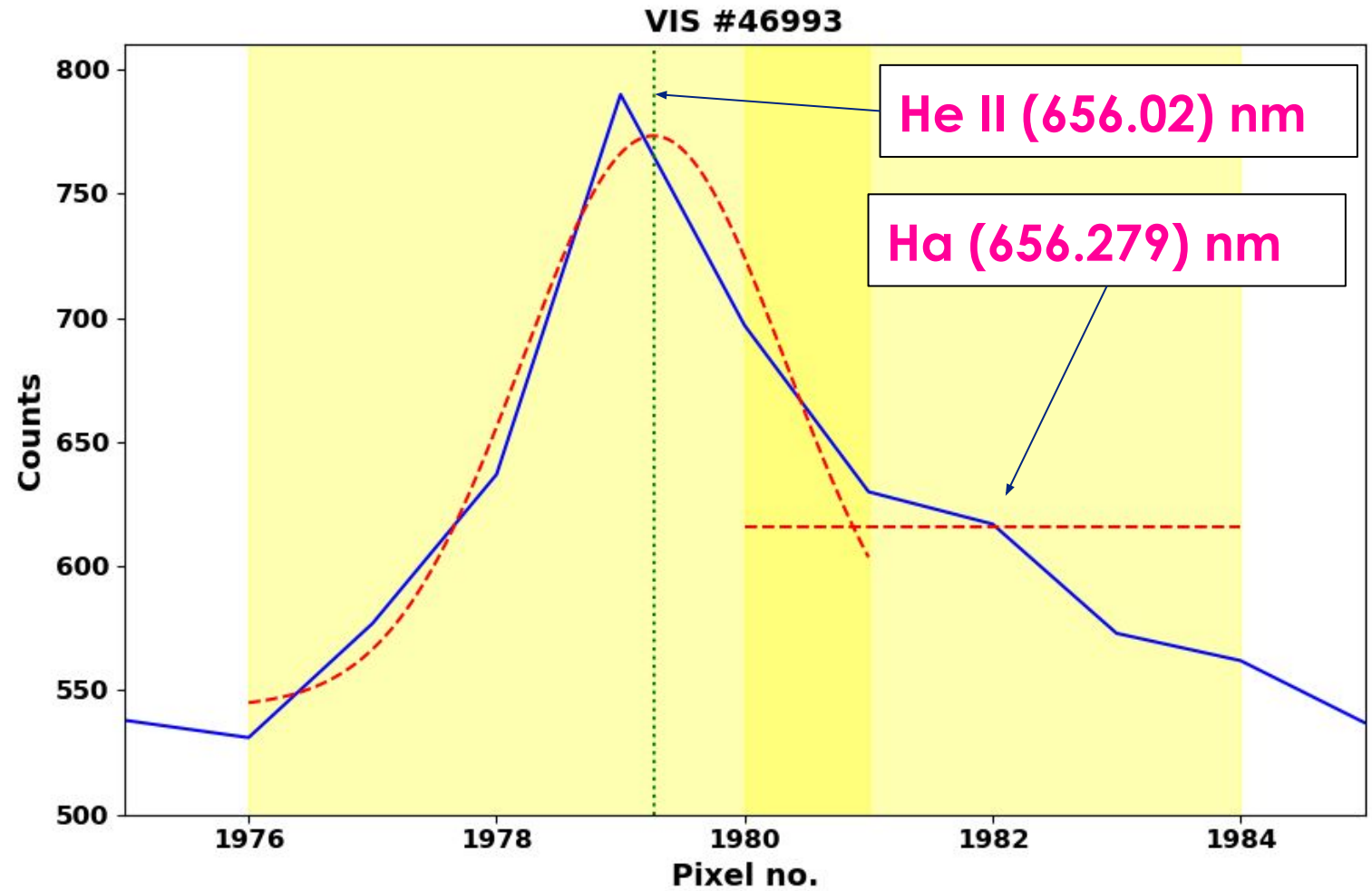
- [1] P. Macha et al 2023 Nucl. Fusion 63 104003; <https://doi.org/10.1088/1741-4326/acf1af>
- [2] Kramida, A., Ralchenko, Yu., Reader, J. and NIST ASD Team (2024). NIST Atomic Spectra Database (version 5.12), [Online]. Available: <https://physics.nist.gov/asd> [Mon Dec 09 2024]. National Institute of Standards and Technology, Gaithersburg, MD. DOI: <https://doi.org/10.18434/T4W30F>
- [3] GUIRLET Rémy, providing graph of the fractional abundances of Carbon
- [4] Finken et al. (2007) Improved Confinement due to Open Ergodic Field Lines Imposed by the Dynamic Ergodic Divertor in TEXTOR. Physical review letters. 98. 065001. 10.1103/PHYSREVLETT.98.065001.

Extra slides

Big peak →
Fit of all peaks
not possible

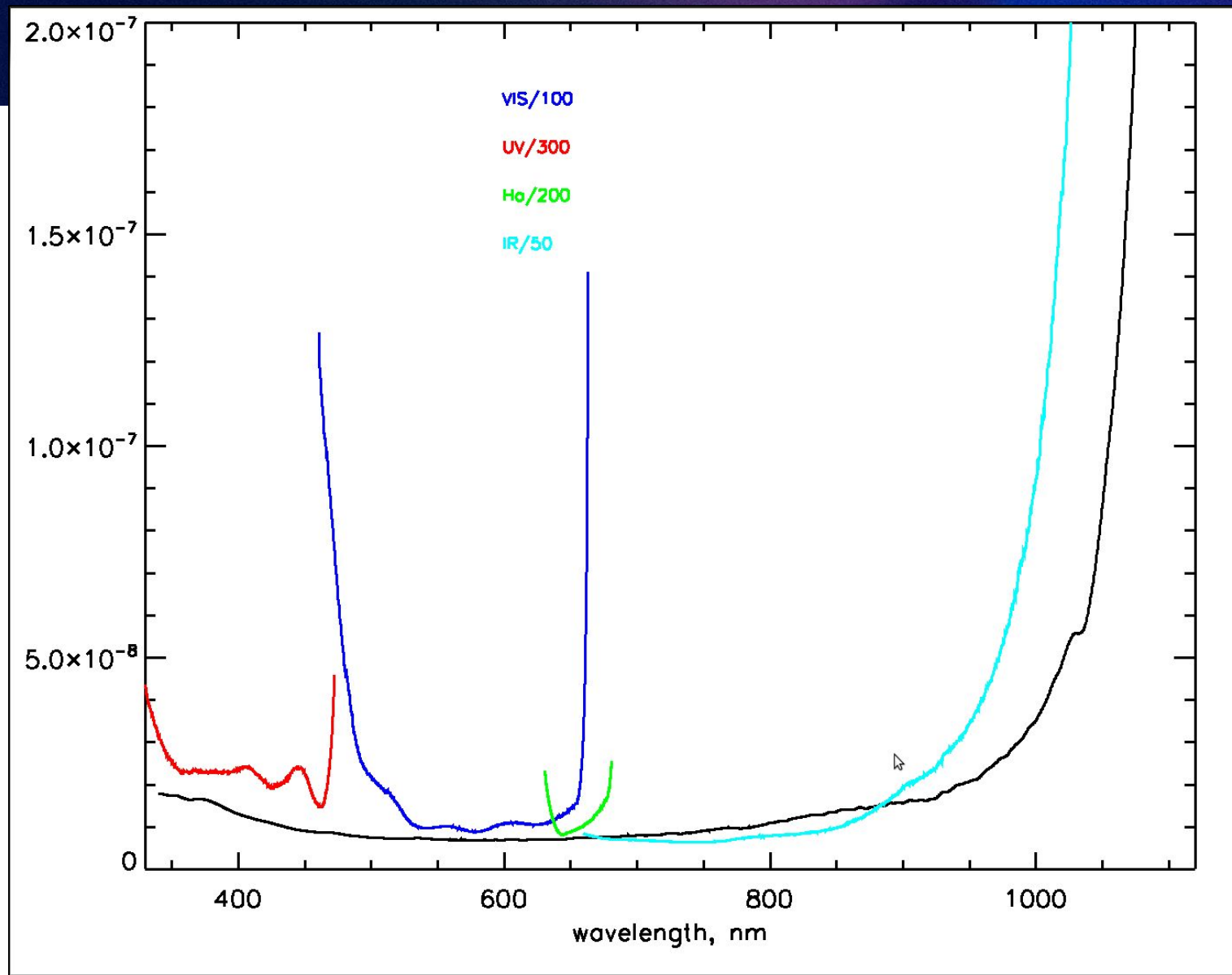


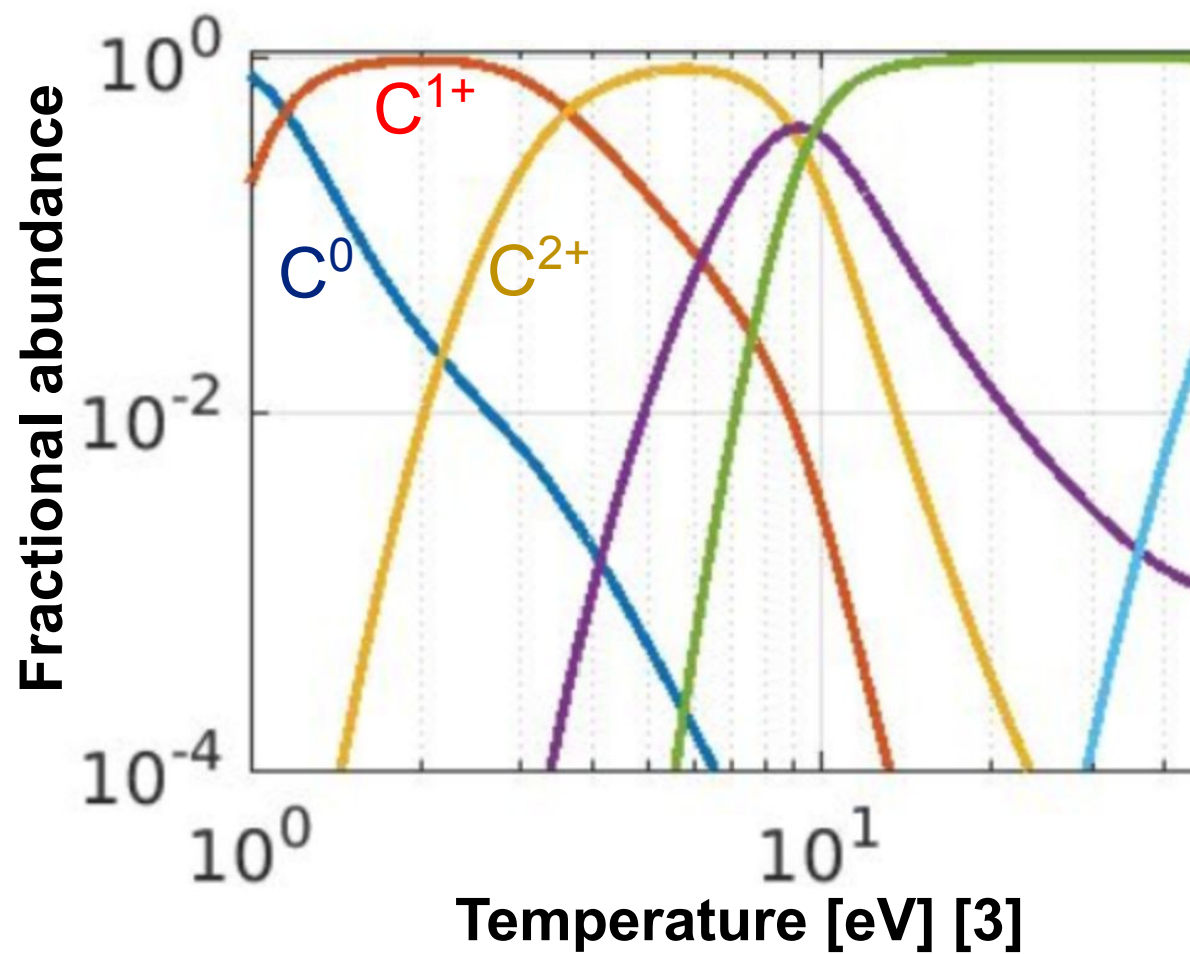
He II & Ha peak →
Fit of both peaks
not possible



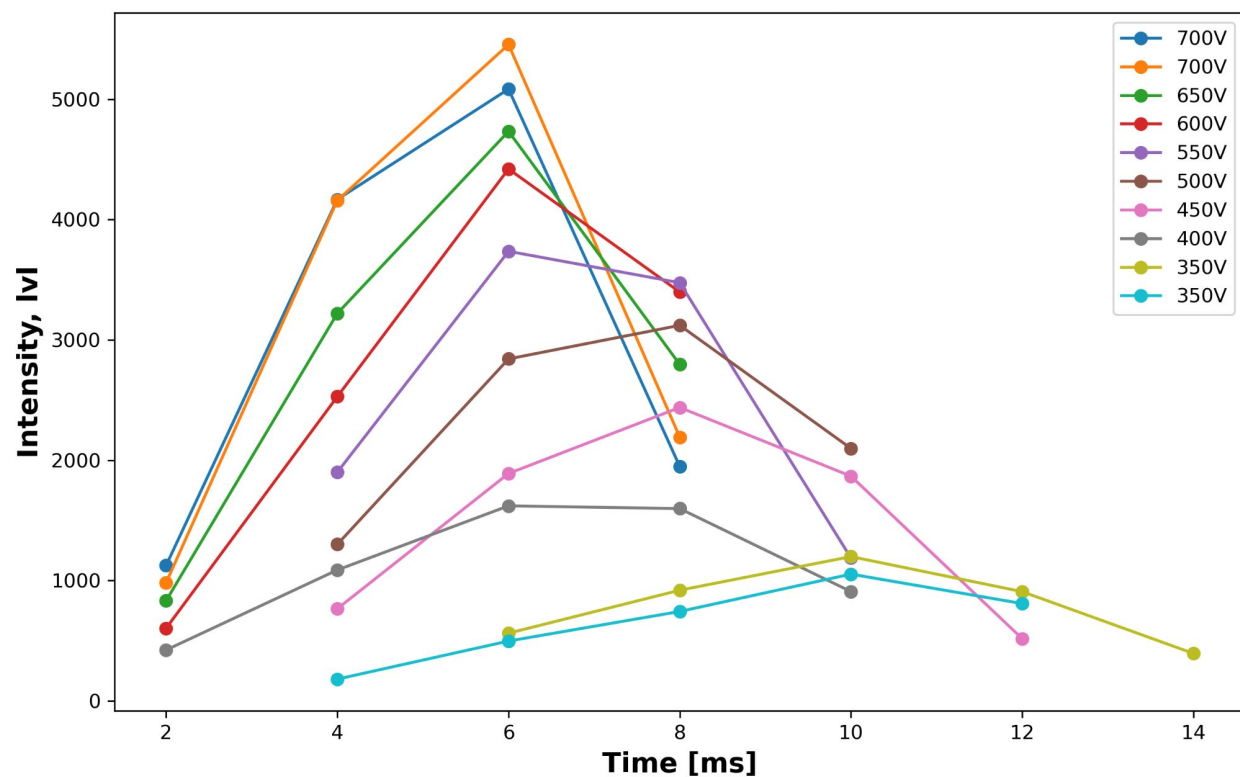
Absolute Calibration Constant (1/QE)

- VIS/100
- UV/300
- H α /200
- IR/50

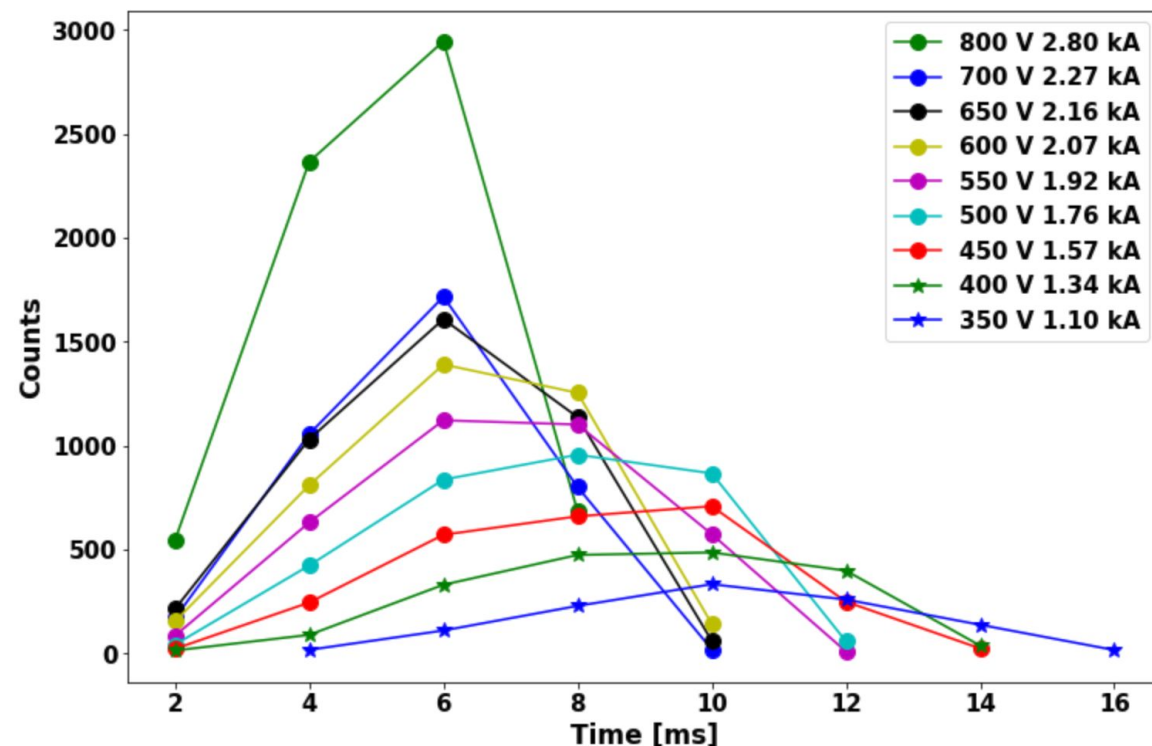




Big Peak Time Evolution (IRVISUV)

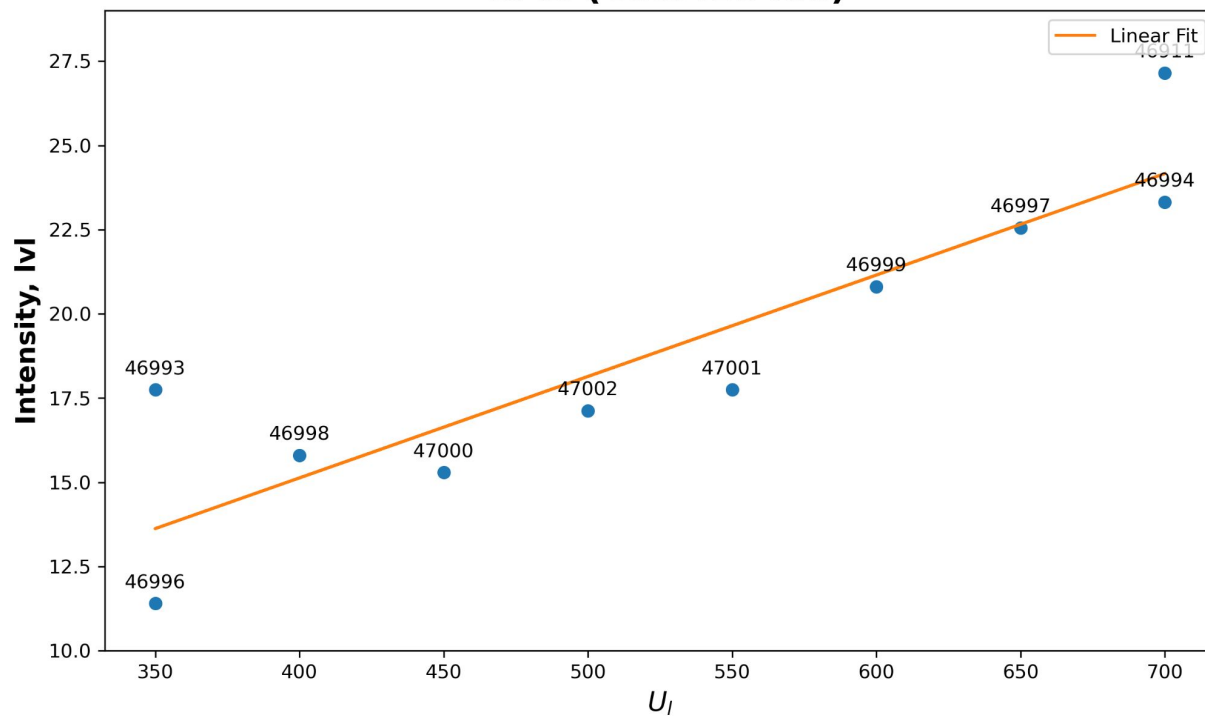


C II Time Evolution (IRVISUV)

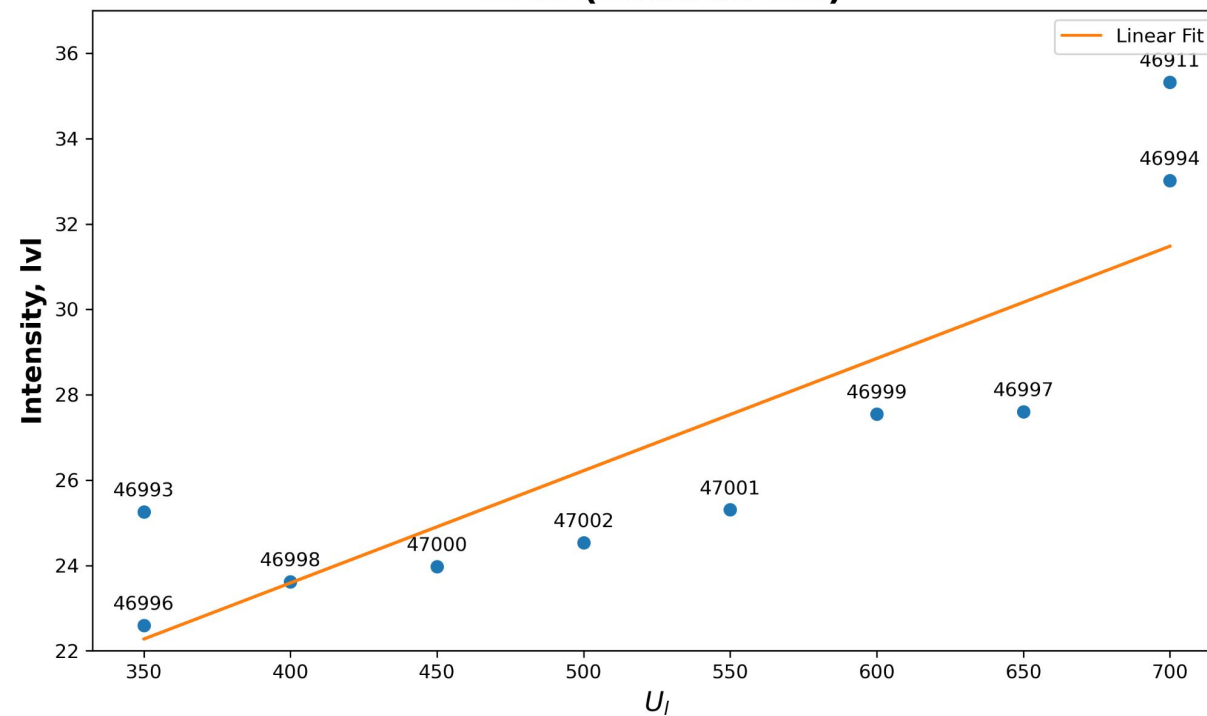


Cl III doublet: 460.425, 460.768 nm

UV Spectrometer:
Cl III (460.425 nm)

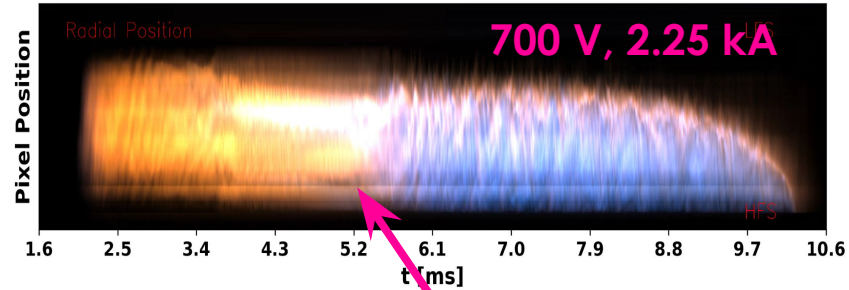


UV Spectrometer:
Cl III (460.768 nm)

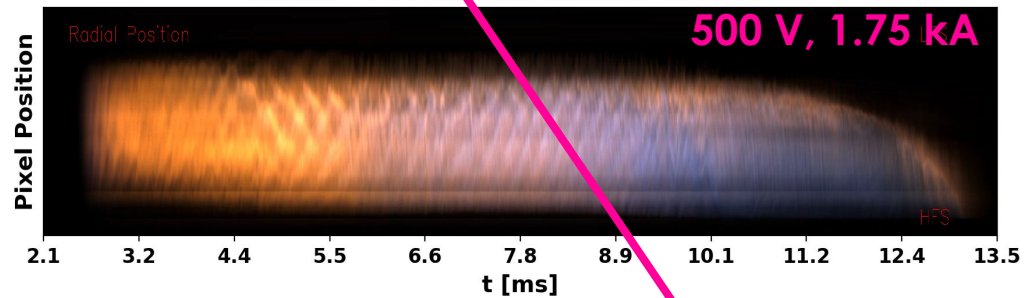




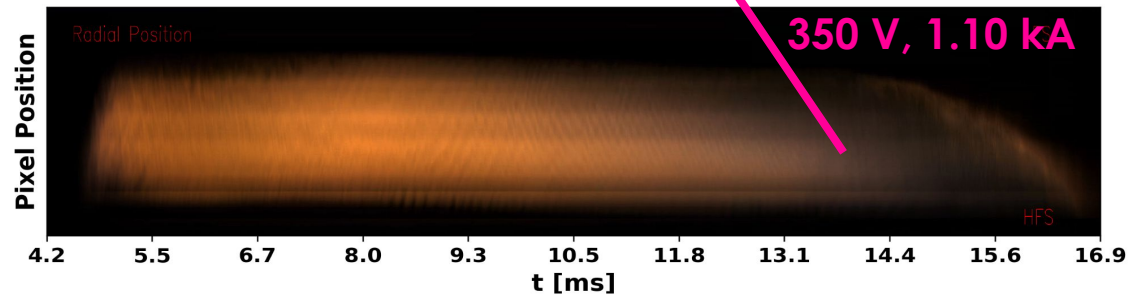
Plasma overview from cameras in #46994



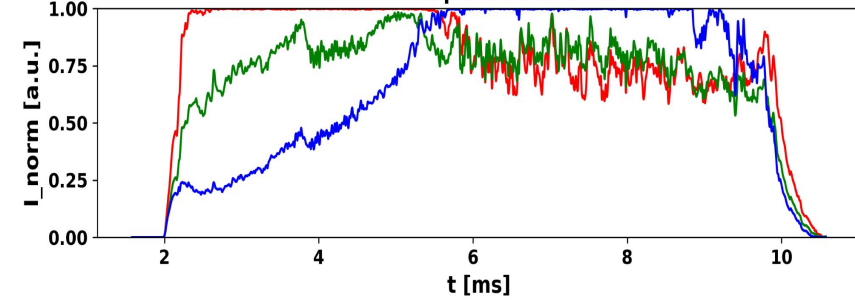
Plasma overview from cameras in #47002



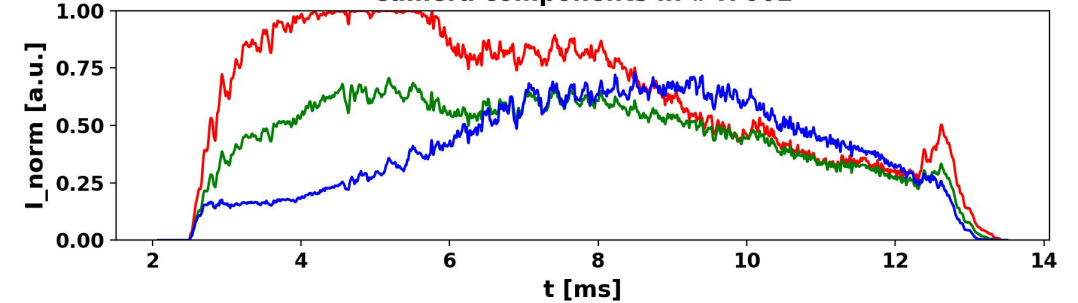
Plasma overview from cameras in #46996



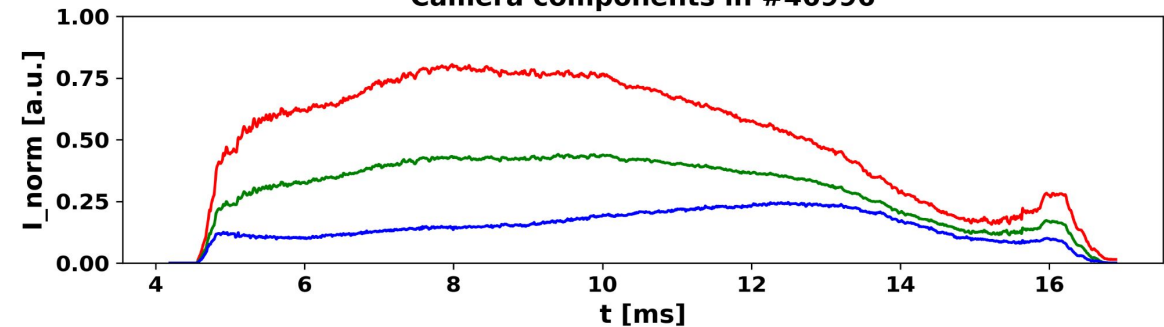
Camera components in #46994



Camera components in #47002

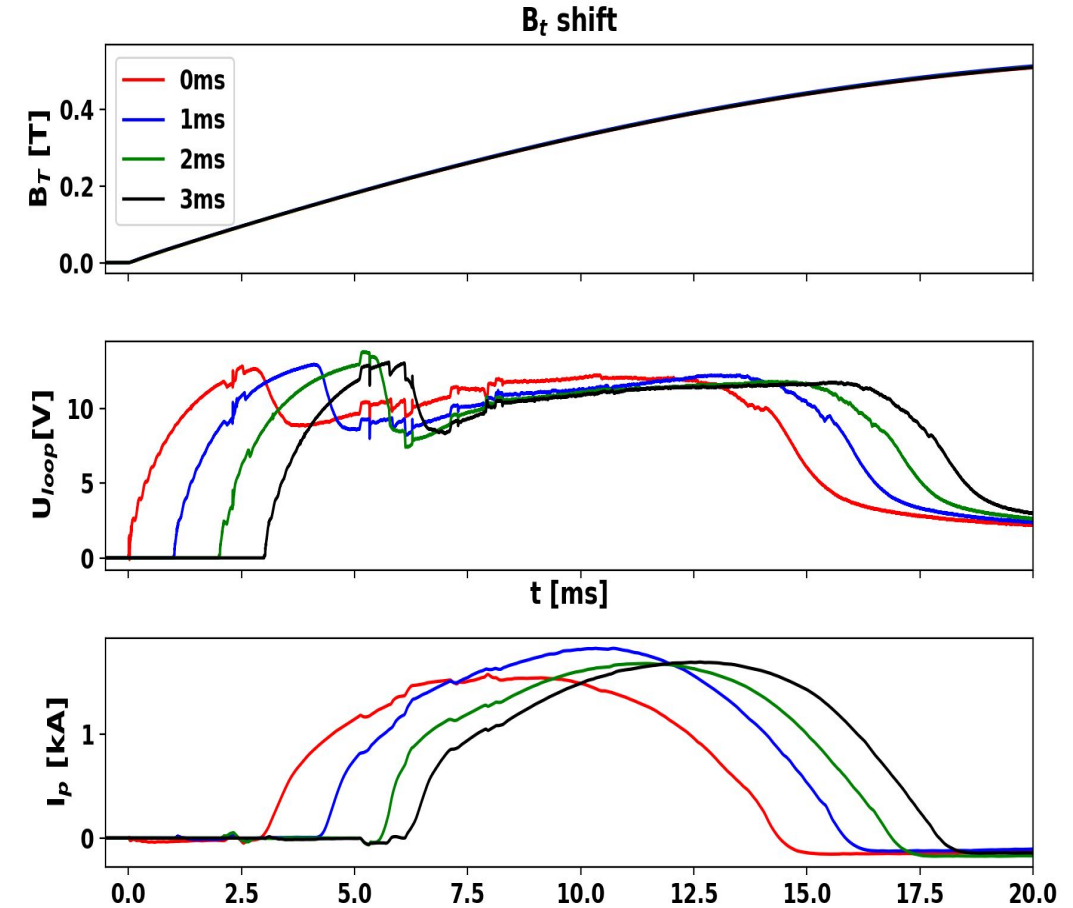
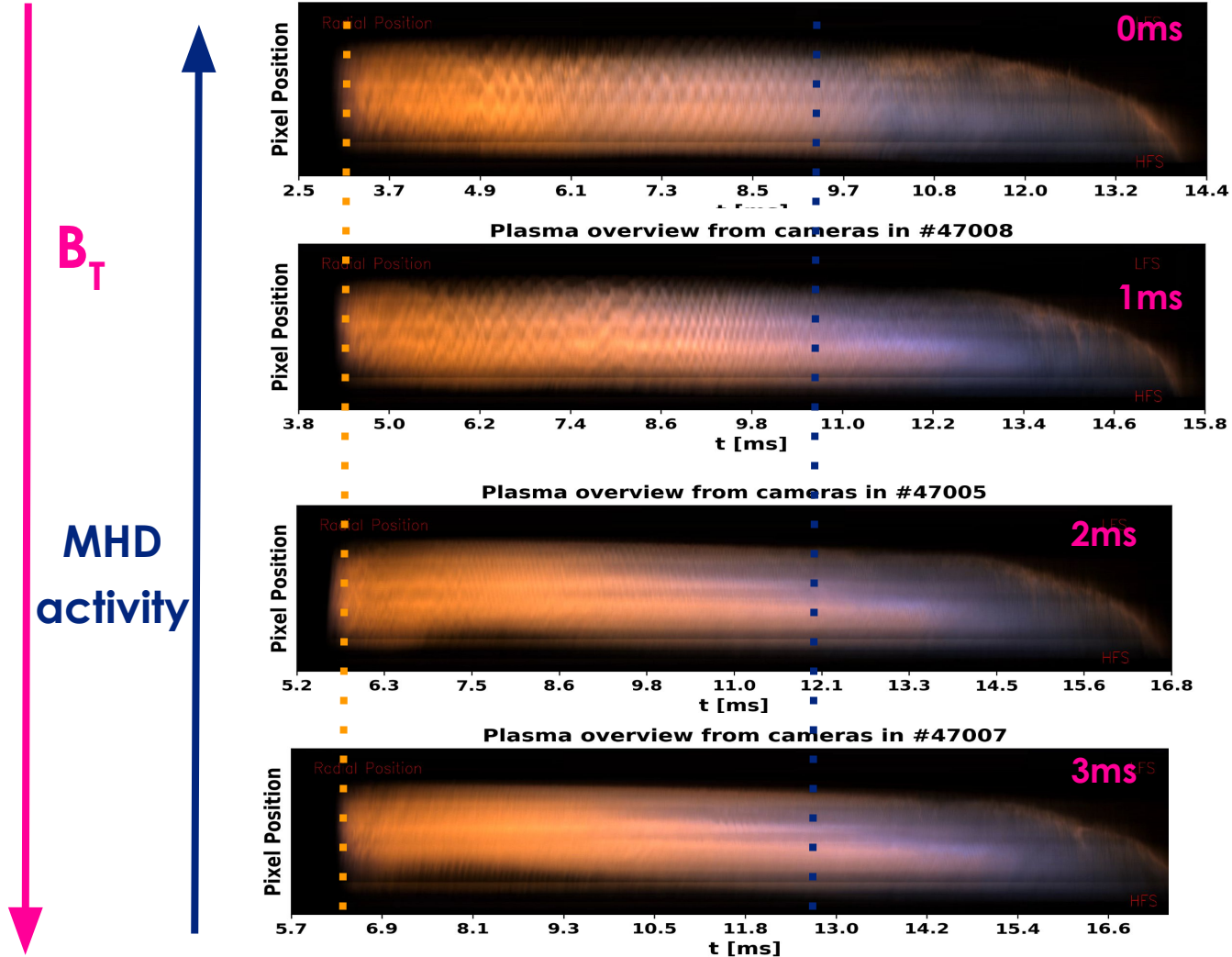


Camera components in #46996



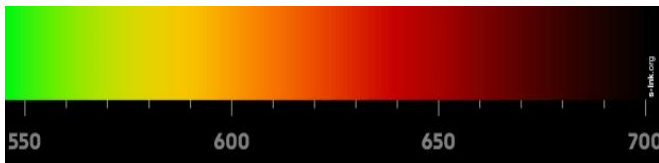
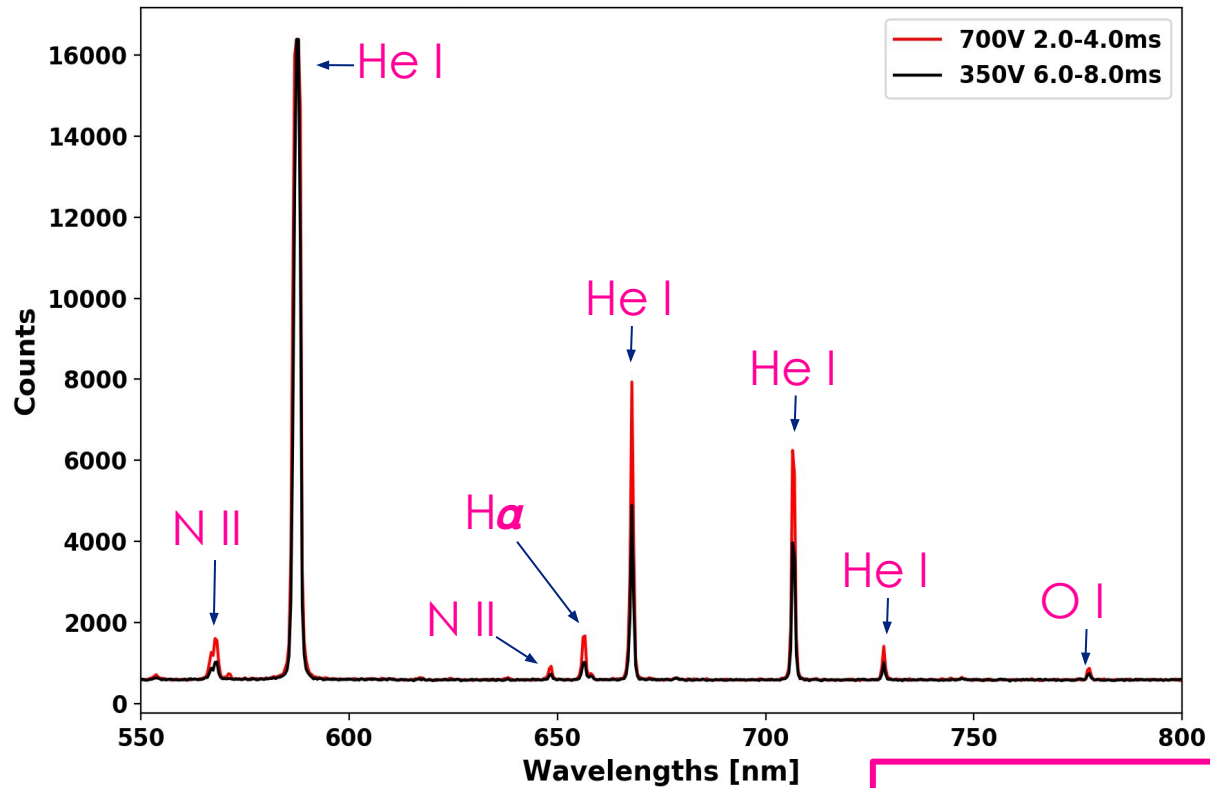


B_T shift scan: Color transition, MHD



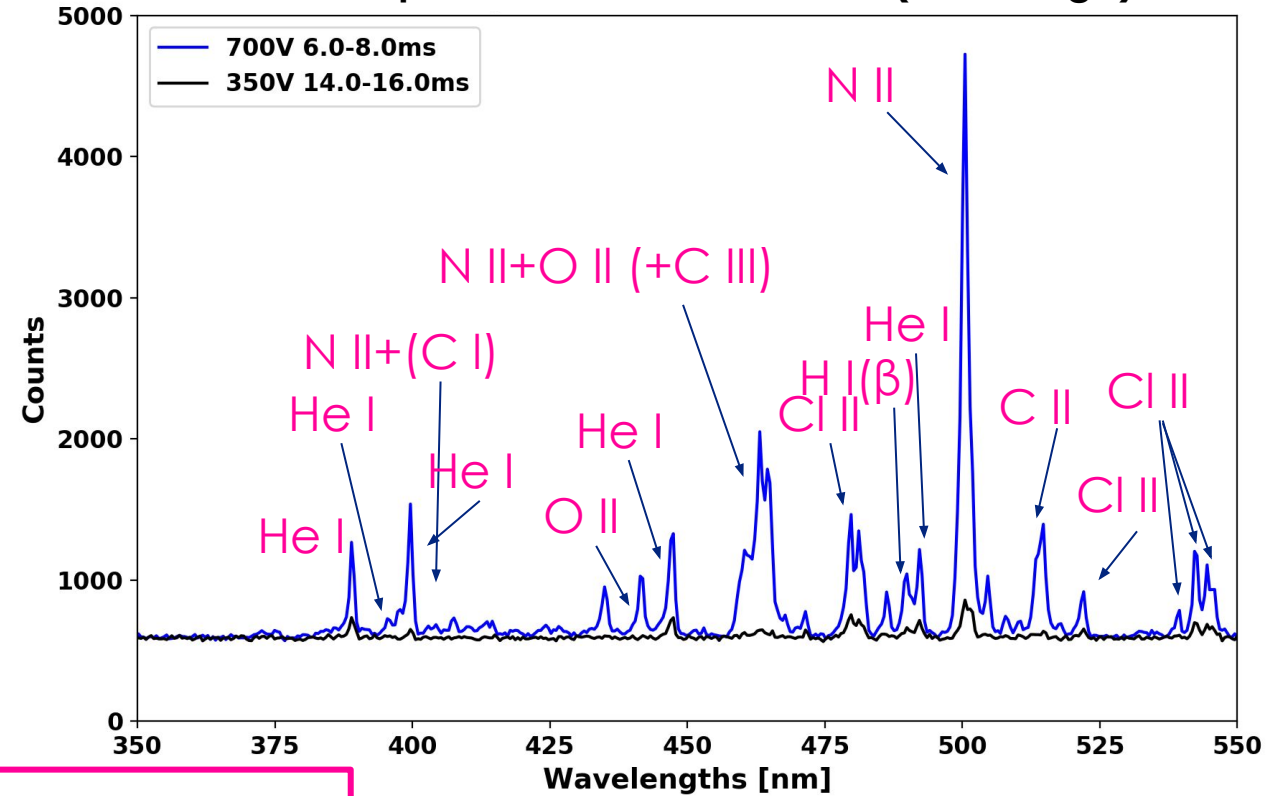


Spectra before the transition (Red range)



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Spectra after the transition (Blue range)



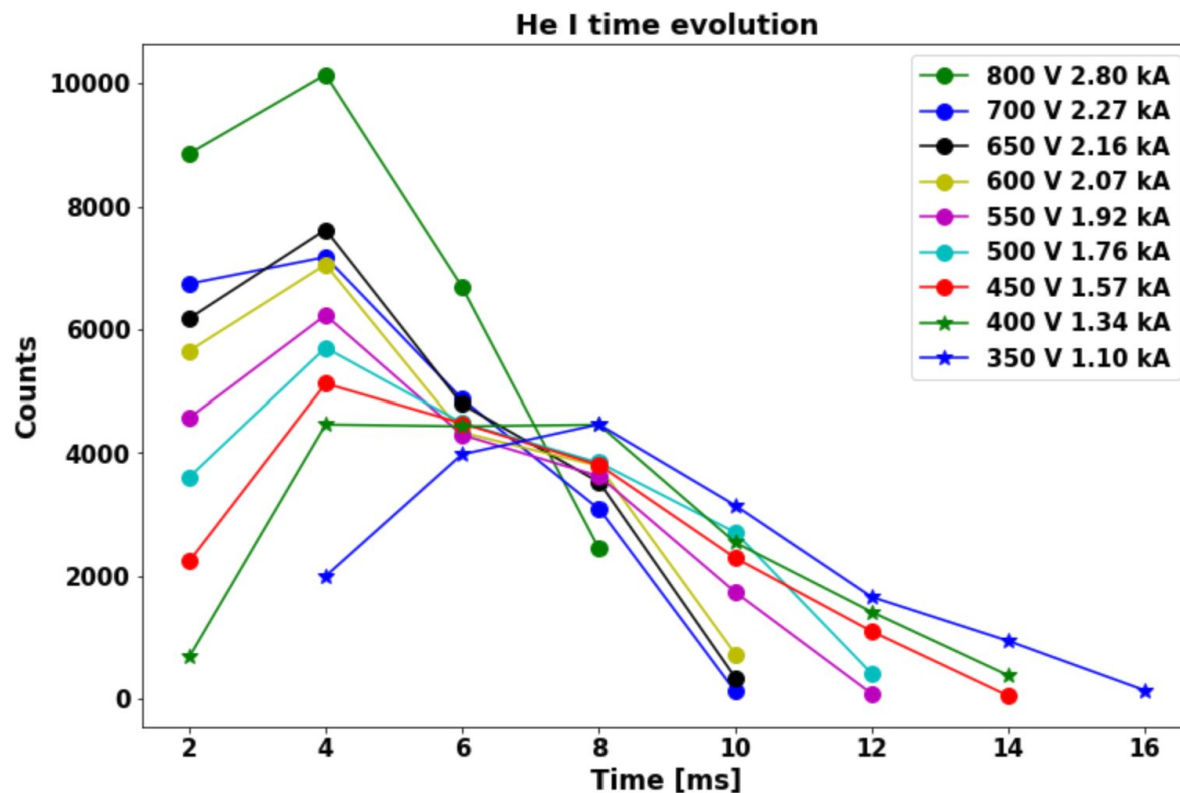
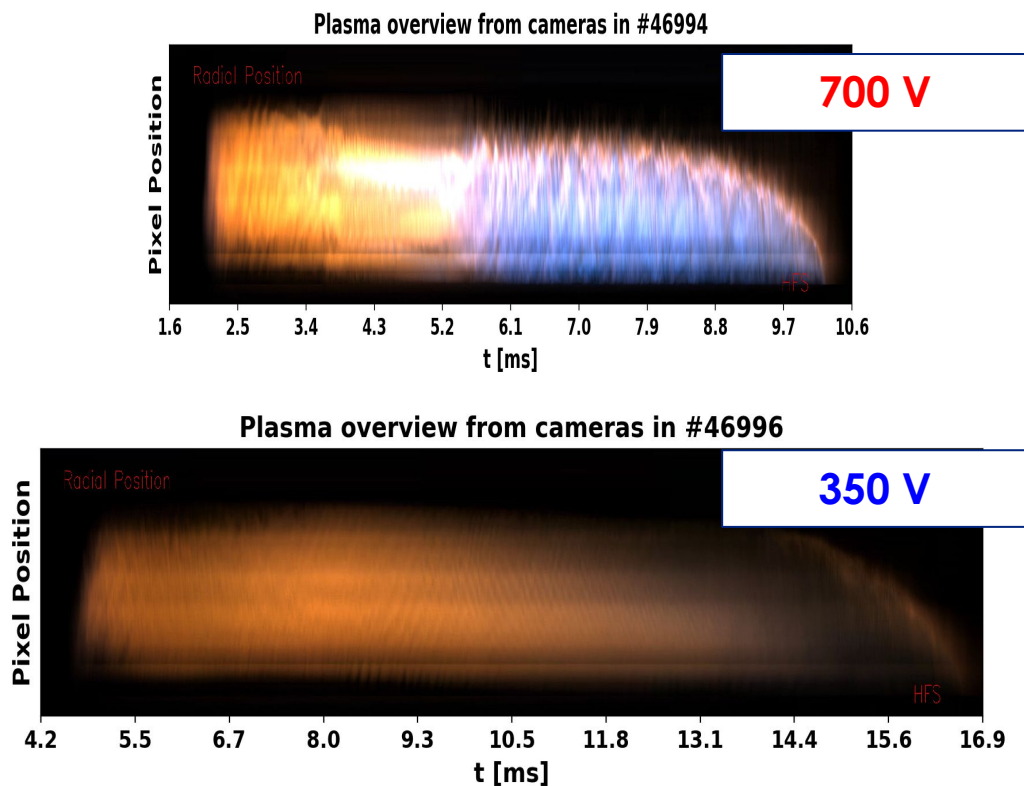
<https://creativecommons.org/licenses/by-sa/4.0/>

Working gas: He I
Impurities:
H I, N II, C II, Cl II, O I, O II
59 lines, 2ms exposure time



He I (667.8 nm)

$$\text{Intensity} \sim n_e * n_{\text{ion}} * C_{\text{ex}}(n_e, T_e)$$

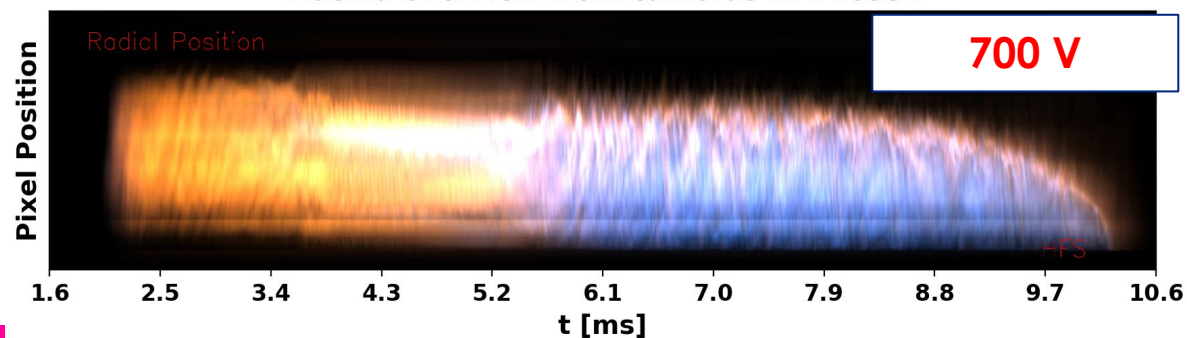


Source: Working gas

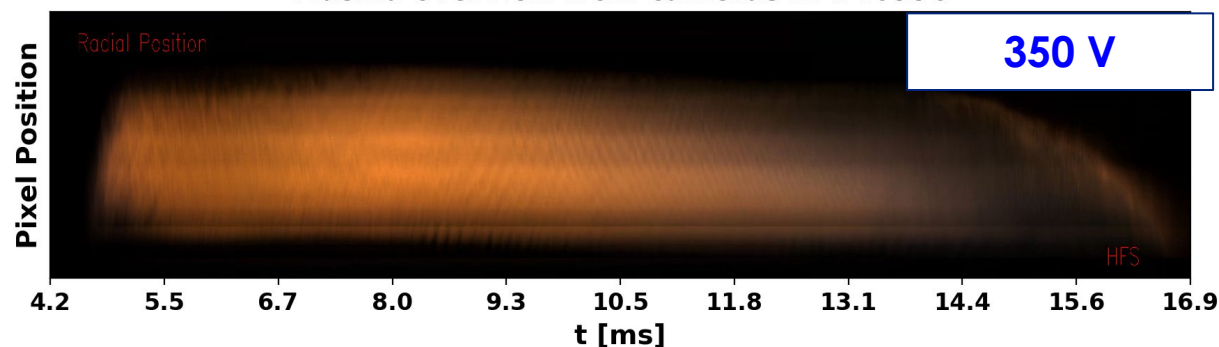


H α (656.3 nm)

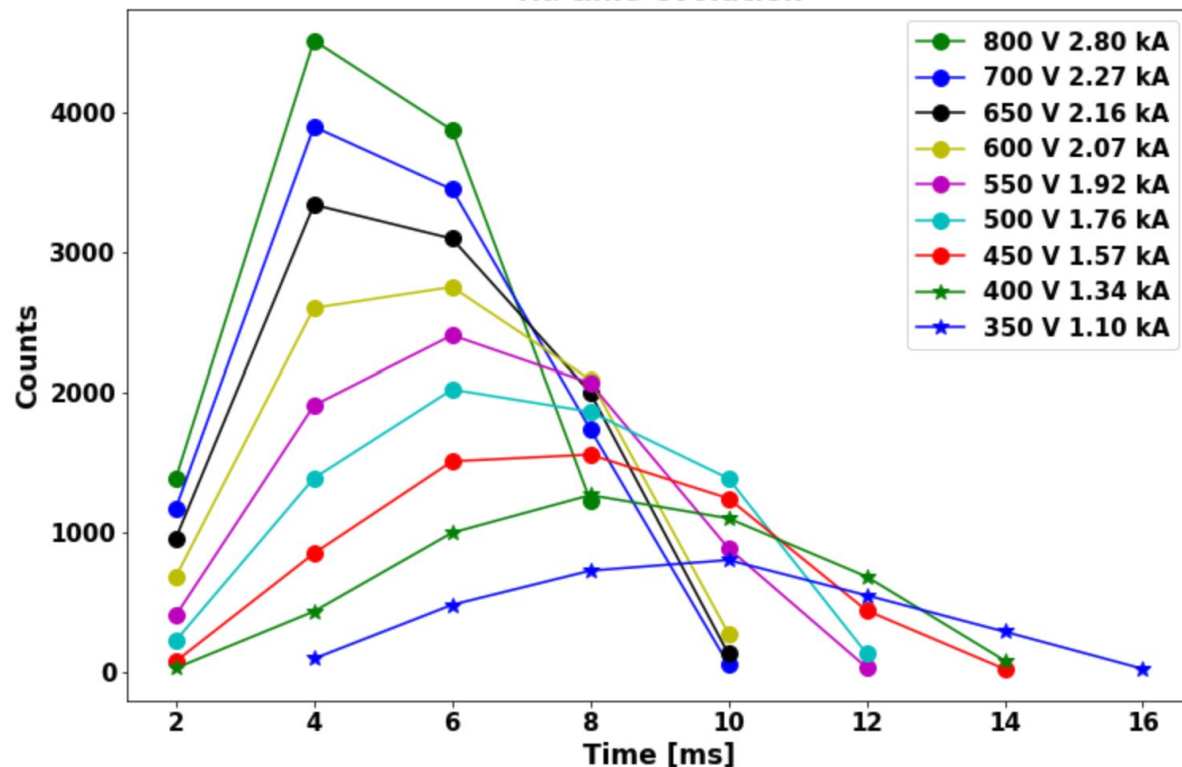
Plasma overview from cameras in #46994



Plasma overview from cameras in #46996



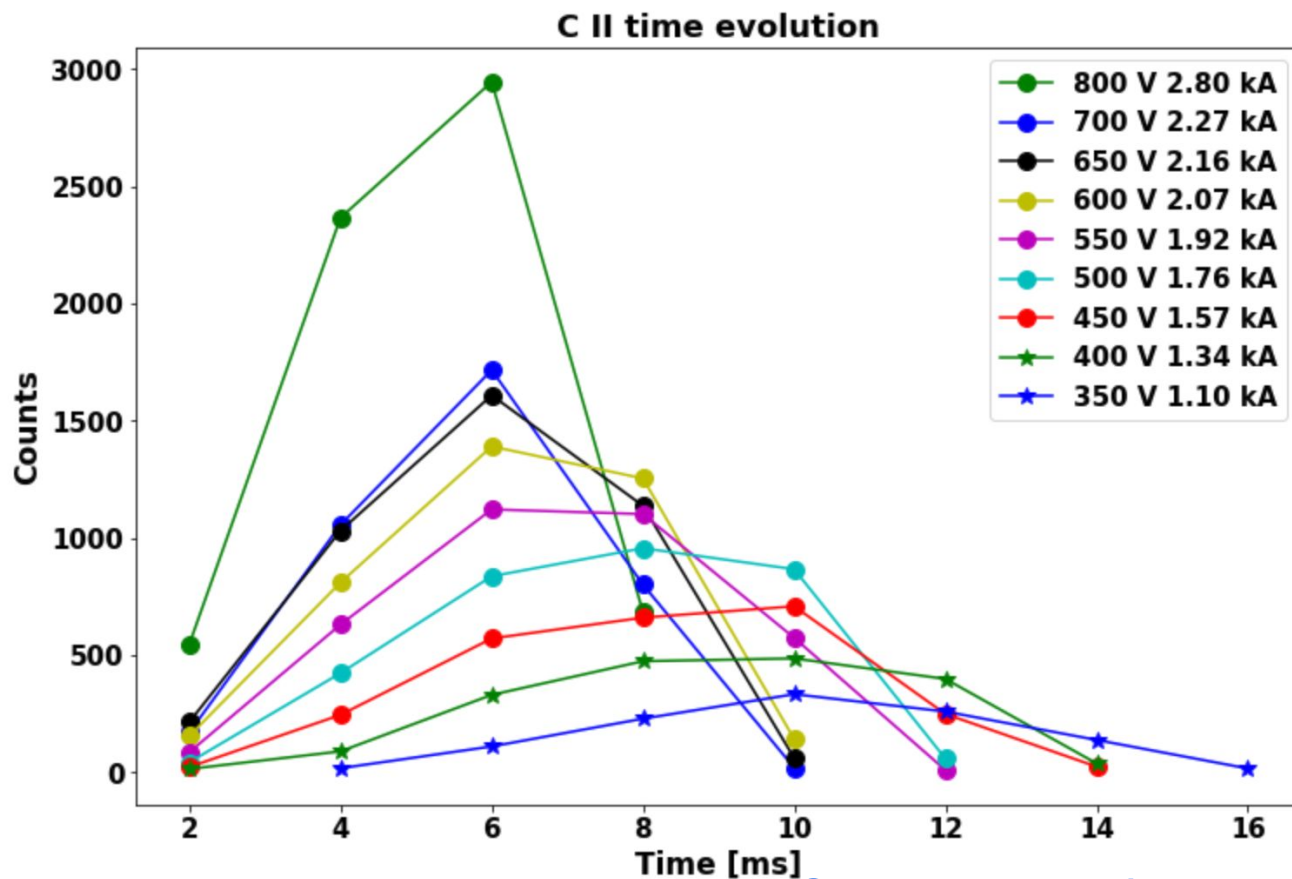
H α time evolution



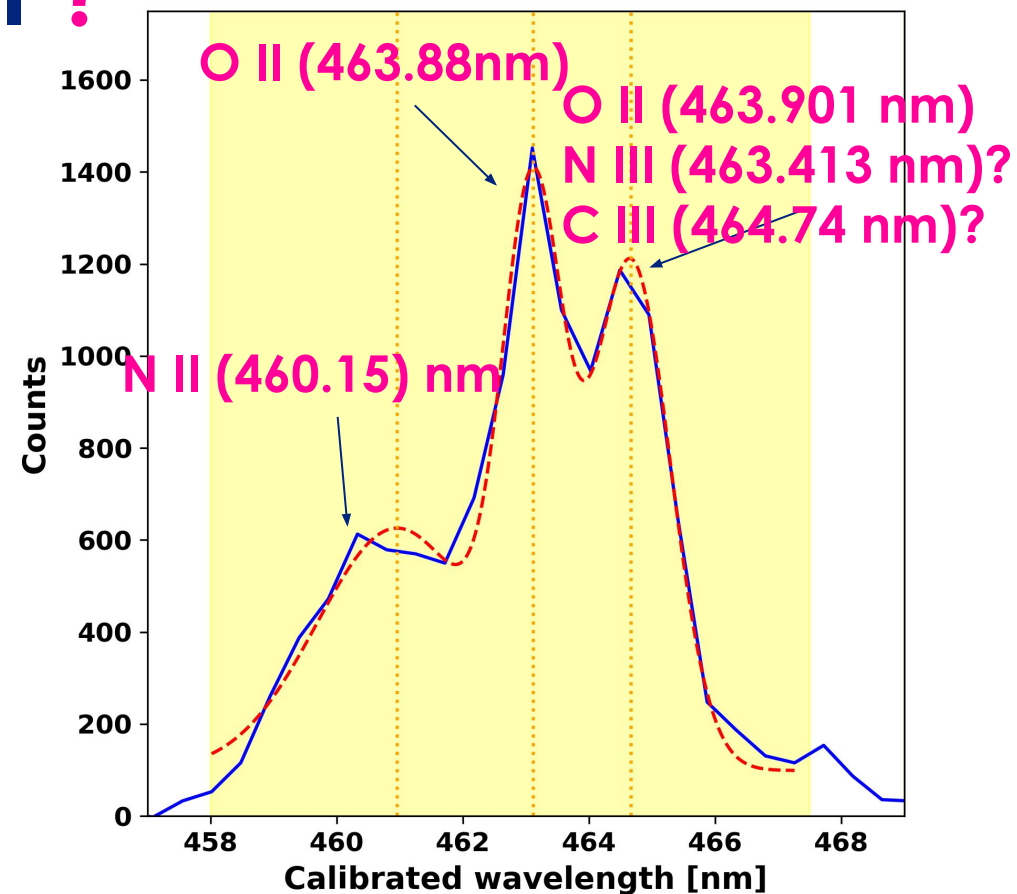
Source: Atmosphere (H₂O), in-vessel materials, previous shots, PVC



C II (657.80nm)

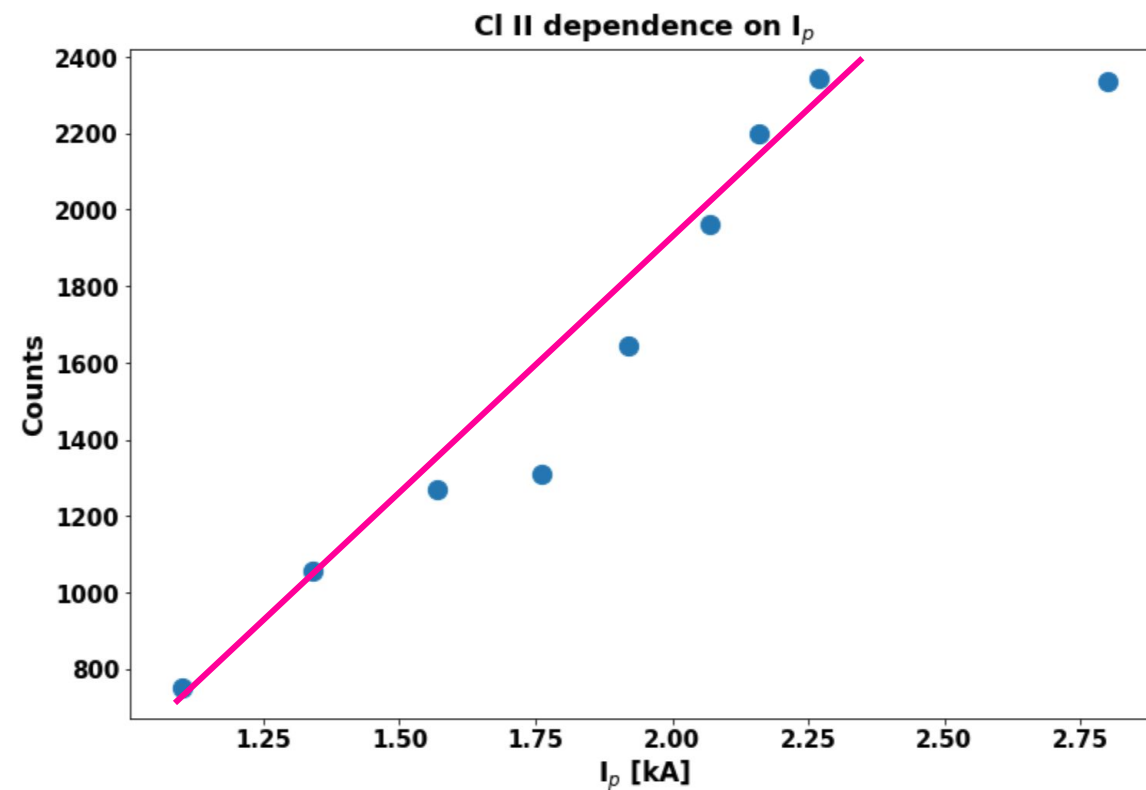
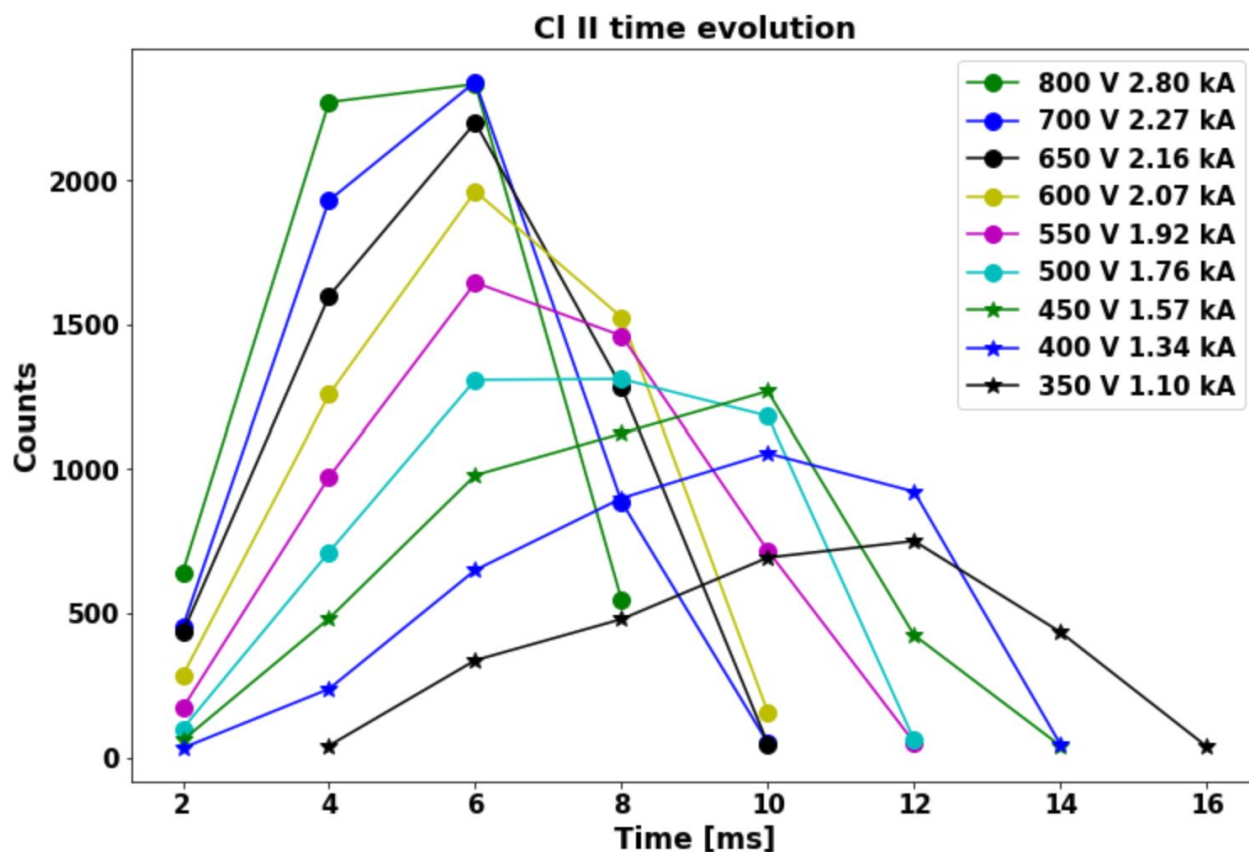


C III ?



Source: Atmosphere (CO₂), PVC, Steel, pumping

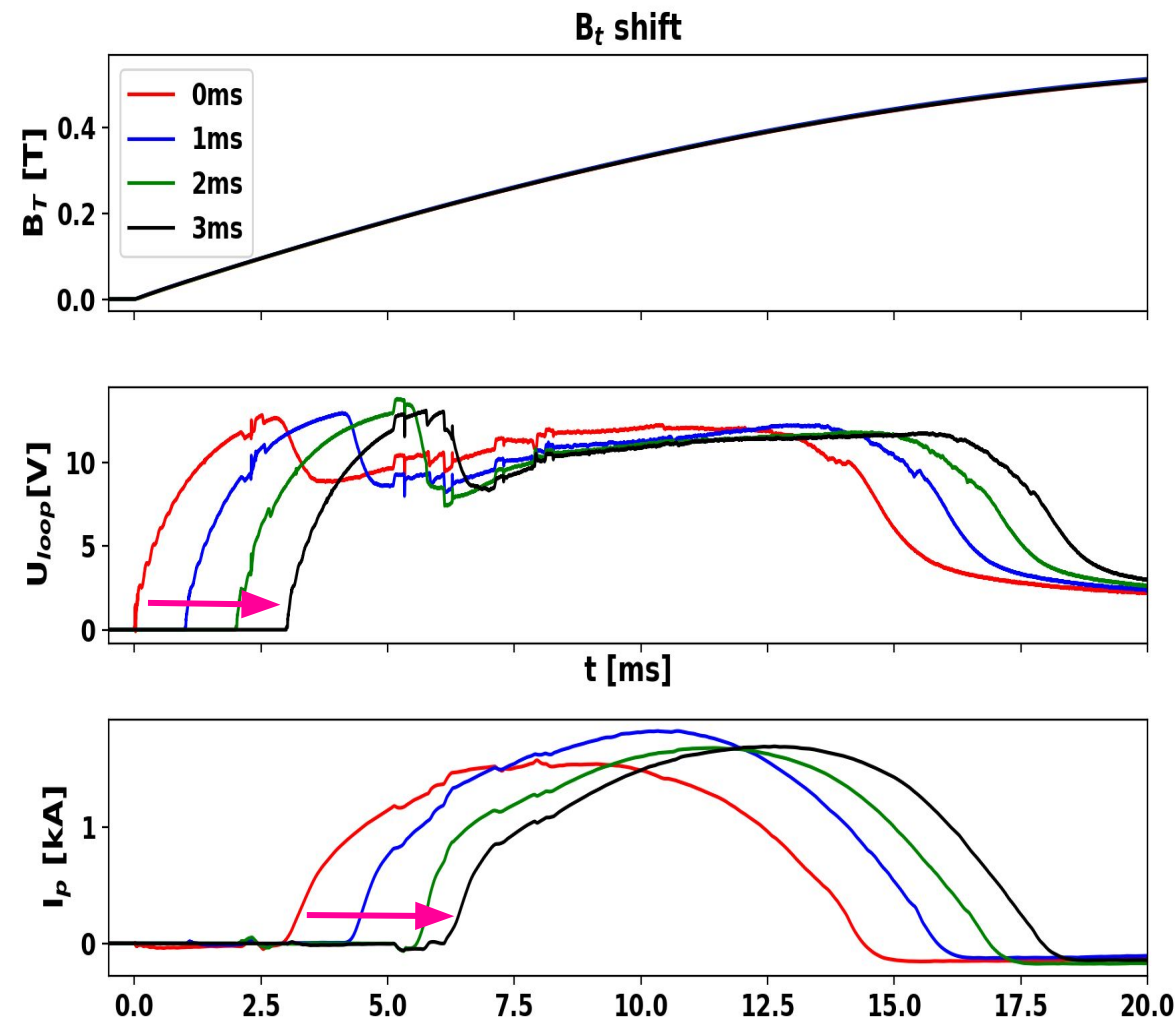
Cl II (479.46 nm)



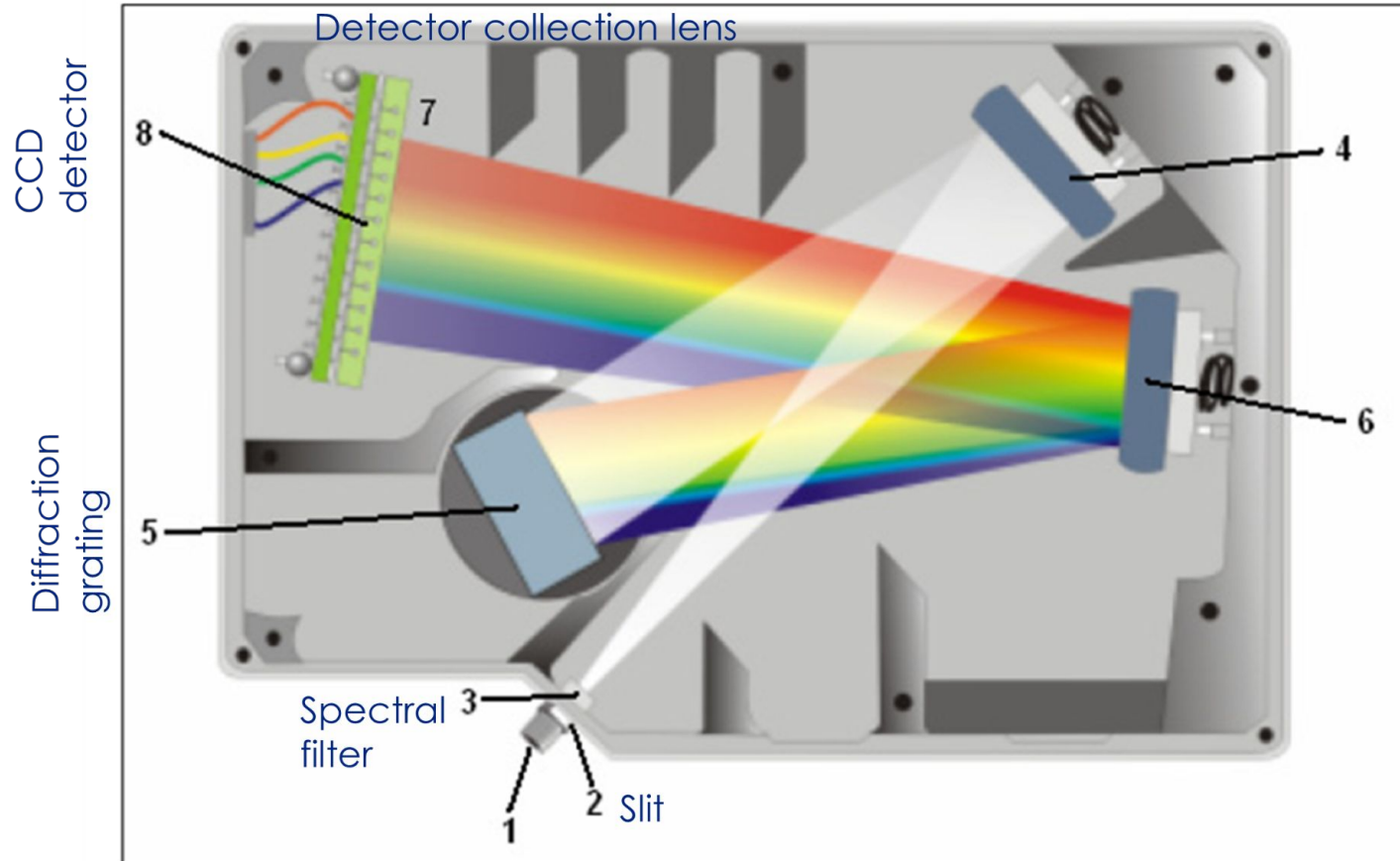
Source: PVC → cable insulation!

SHOT NUMBER	time delay [ms]	U (Current drive) [V]	$I_{P, MAX}$ [kA]	B_t (at $I_{P, MAX}$) [T]
47000	0	450	1.6	0.27
47008	1	450	1.8	0.35
47005	2	450	1.7	0.37
47007	3	450	1.7	0.39

Table of shots with similar current drive values but different toroidal magnetic field shift.

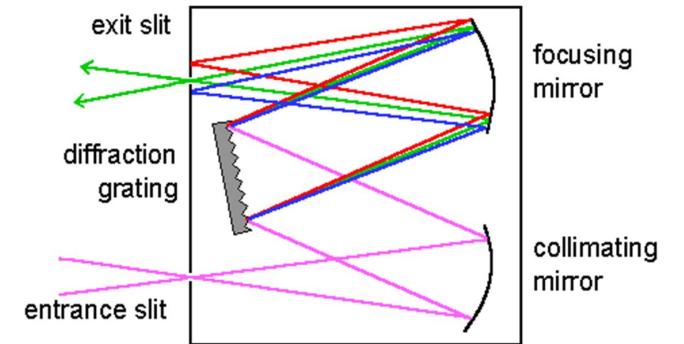


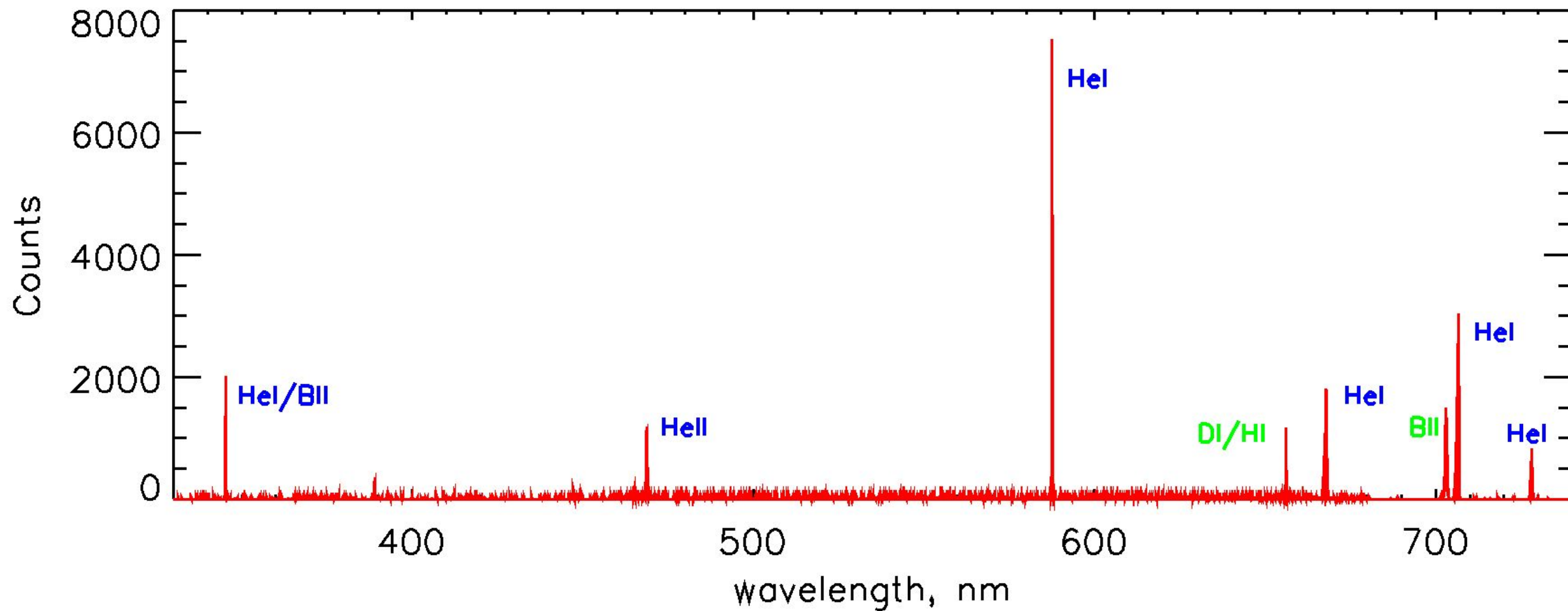
HR2000+: crossed **Czerny-Turner spectrometer**
observing plasma using optical fibre



SMA connector for optical fiber

Czerny-Turner set-up



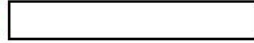


IR 20 μ s, VIS 3 μ s, UV 50 μ s, Halpha 20 μ s

wavelength - color

<https://www.agcled.com/blog/basic-information-of-the-visible-light.html>

<https://www.pythoninformer.com/computer-science/colour/rgb-colour/>

		
rgb(0%, 0%, 0%)	rgb(50%, 0%, 0%)	rgb(100%, 0%, 0%)
		
rgb(0%, 50%, 0%)	rgb(50%, 50%, 0%)	rgb(100%, 50%, 0%)
		
rgb(0%, 100%, 0%)	rgb(50%, 100%, 0%)	rgb(100%, 100%, 0%)
		
rgb(0%, 0%, 50%)	rgb(50%, 0%, 50%)	rgb(100%, 0%, 50%)
		
rgb(0%, 50%, 50%)	rgb(50%, 50%, 50%)	rgb(100%, 50%, 50%)
		
rgb(0%, 100%, 50%)	rgb(50%, 100%, 50%)	rgb(100%, 100%, 50%)
		
rgb(0%, 0%, 100%)	rgb(50%, 0%, 100%)	rgb(100%, 0%, 100%)
		
rgb(0%, 50%, 100%)	rgb(50%, 50%, 100%)	rgb(100%, 50%, 100%)
		
rgb(0%, 100%, 100%)	rgb(50%, 100%, 100%)	rgb(100%, 100%, 100%)

380 V 450 B 495 G 570 590 630 Y O 750 R			
Color	Wavelength (nm)	Frequency (THz)	Photon energy (eV)
violet	380–450	670–790	2.75–3.26
blue	450–485	620–670	2.56–2.75
cyan	485–500	600–620	2.48–2.56
green	500–565	530–600	2.19–2.48
yellow	565–590	510–530	2.10–2.19
orange	590–625	480–510	1.98–2.10
red	625–750	400–480	1.65–1.98